

1 **High-MgO lavas associated to CFB as indicators of plume-related**

2 **thermochemical effects: the case of ultra-titaniferous picrite-basalt from**

3 **the Northern Ethiopian-Yemeni Plateau**

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11 **Abstract**

12 A comprehensive petrological and geochemical dataset is reported in order to define the
13 thermo-compositional characteristics of Ti (Fe)-enriched picrite-basalt lavas (HT2, TiO₂ 3-7
14 wt%), erupted close to the axial zone of the inferred Afar mantle plume, at the centre of the
15 originally continuous Ethiopian-Yemeni CFB plateau (ca. 30Ma) which is zonally arranged
16 with progressively lower Ti basalts (HT1, TiO₂ 2-4 wt%; LT, TiO₂ 1-3 wt%) toward the
17 periphery. Integrated petrogenetic modelling based on major and trace element analyses of
18 bulk rocks, minerals and melt inclusions in olivines, as well as Sr-Nd-Pb-He-O isotope
19 compositional variations enables us to make several conclusions. 1) The phase equilibria
20 constraints indicate that HT2 primary picrites were generated at ca. 1570°C mantle potential
21 temperatures (T_p) in the pressure range 4-5 GPa whereas the HT1 and LT primary melts
22 formed at shallower level (<2 to 3 GPa, T_p 1530 °C for HT1 and 1430°C for LT). Thus the
23 Afar plume head was a thermally and compositionally zoned melting region with maximum
24 excess temperatures of 300-350°C with respect to the ambient mantle. 2) The HT2 primary
25 melts upwelled nearly adiabatically to the base of the continental crust (ca. 1 GPa) where

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fractionation of olivine, followed by clinopyroxene, led to variably differentiated picritic and basaltic magmas. 3) Trace element modelling requires that the primary HT2 melts were generated - either by fractional or batch melting (F 9-10%) - from a mixed garnet peridotite source (85%) with 15% eclogite (derived from transitional MORB protoliths included in Panafrican terranes) that has to be considered a specific Ti-Fe and incompatible element enriched component entrained by the Afar plume. 4) The LT, HT1 and HT2 lavas have $^{143}\text{Nd}/^{144}\text{Nd} = 0.5131\text{-}0.5128$, whereas Sr-Pb isotopes are positively correlated with TiO_2 , varying from $^{87}\text{Sr}/^{86}\text{Sr} 0.7032$ and $^{206}\text{Pb}/^{204}\text{Pb} 18.2$ in LT basalts to $^{87}\text{Sr}/^{86}\text{Sr} 0.7044$ and $^{206}\text{Pb}/^{204}\text{Pb} 19.4$ in HT2 picrite-basalts. High $^3\text{He}/^4\text{He}$ (15-20 R_A) ratios are exclusively observed in HT2 lavas, confirming earlier evidence that these magmas require a component of deep mantle in addition to eclogite, while the LT basalts may more effectively reflect the signature of the pre-existing mantle domains. The comparison between high-MgO (13-22%) lavas from several Phanerozoic CFB provinces (Karoo, Paraná-Etendeka, Emeishan, Siberia, Deccan, North Atlantic Province) shows that they share extremely high mantle potential temperatures (T_p 1550-1700°C) supporting the view that hot mantle plumes are favoured candidates for triggering many LIPs. However, the high incompatible element and isotopic variability of these high-MgO lavas (and associated CFB) suggest that plume thermal anomalies are not necessarily accompanied by significant and specific chemical effects, which depend on the nature of mantle materials recycled during the plume rise, as well as by the extent of related mantle enrichments (if any) on the pre-existing lithospheric section.

1. Introduction

There is a longstanding debate on the generation of Large Igneous Province (LIP), either attributed to mantle plumes, melting anomalies in compositionally heterogeneous mantle (see Special Papers by Ernst and Buckan, 2001a; Foulger et al., 2005, Foulger and Jurdy, 2007; Beccaluva et al., 2011a) or even to mantle warming beneath continents (Coltice et al., 2009).

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52 A great contribution to the plume debate has been provided by the tracking through the
53 geological time (from Archean to Phanerozoic) of large amounts of erupted magmas
54 emplaced in a short time in both continental and oceanic settings (Ernst and Buckan, 2001b;
55 Courtillot et al., 2003; Safonova and Santosh, 2014; Reichow et al., 2009). Among the
56 various magma types observed in LIPs, an increasing interest has been devoted to high-Mg
57 lavas that—owing to their primitiveness, deep-seated origin and general lack of crustal
58 contamination could be considered effective proxies for plume-related magmas.
59 High-Mg lavas such as picritic basalts, (ferro)picrites and meimechites, are observed in many
60 Phanerozoic LIP where they represent minor but significant components in both continental
61 and oceanic flood basalt sequences (Karoo, Deccan, Siberia, Ethiopia-Yemen, Paraná-
62 Etendeka, North Atlantic Province-NAP, Emeishan and Caribbean). The intrinsic
63 characteristics of high-Mg magmas led many authors to consider them as the deepest and
64 hottest melting products of an upwelling mantle plume (Ellam and Cox, 1991; Ellam et al.,
65 1992; Thomson et al., 2001; Herzberg and O’Hara, 2002; Campbell and Davis, 2006; Carlson
66 et al., 2006; Beccaluva et al., 2009; Zhang et al., 2010), thus more effectively recording its
67 thermo-chemical effects with respect to the associated basalts. Accordingly, some
68 compositional features of these high-Mg magmas -such as low-Al, high-Fe, variable Ti and
69 Ocean Island Basalt (OIB)-like trace element and isotopic signatures- have been considered
70 indicative of generation at high P-T conditions with the involvement of pyroxenite/eclogite
71 components in their mantle sources (Gibson et al., 2000; Beccaluva et al., 2009; Sobolev et
72 al., 2007 and 2009; Buslov et al., 2010; Kamenetskey et al., 2012; Heinonen et al., 2013 and
73 2014).

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74 Application of olivine-melt equilibria constraints in high-Mg lavas, generally allows
75 reliable evaluations of olivine cumulus/fractionation processes as well as reconstruction of
76 primary magmas and the relative mantle potential temperatures (Herzberg et al., 2007). On
77 the other hand, there are several aspects that need additional investigations, in particular, the

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78 significance of picrites with respect to the compositional variability and zonation of the
79 associated basalts (Low-Ti, High-Ti) as well as their mutual spatial/temporal relationships
80 which are often unclear. In general, there is a need for a better understanding of the role of
81 high-Mg magmas in the whole thermo-compositional and spatial development of a LIP and,
82 by implication, the causative events of thermal anomalies in the mantle.

83 In this context, the 30 Ma Northern Ethiopian-Yemeni LIP is of particular interest as it
84 consists of a large originally continuous CFB plateau, characterized by zonal arrangement
85 with picritic-basaltic lavas at the centre corresponding to the Afar triple junction which is
86 considered the axis of a deep mantle plume active beneath the Afro-Arabian region during
87 Oligocene (Beccaluva et al., 2009; Natali et al., 2011; Rooney et al., 2012). This picrite-
88 basalt association includes the most Ti enriched high-Mg lavas so far observed in CFB
89 sequences (TiO_2 up to 7 wt%), comparable only with meimechites from the Siberian Traps
90 (Carlson et al., 2006; Sobolev et al., 2009). Their peculiar composition together with their
91 position at the core of the Oligocene CFB plateau make these rocks extremely interesting to
92 evaluate both the temperature of the Afar mantle plume at the time of CFB eruption and its
93 contribution to the composition of the related. Moreover, systematic investigations on mantle
94 xenoliths included in Neogene alkaline lavas overlying the CFB plateau, and from
95 neighbouring districts, now provide direct insights on the nature of the ambient mantle and
96 allow important constraints to be placed on the primary melt petrogenesis in the Afro-
97 Arabian domain (Beccaluva et al., 2011b; Bianchini et al, 2014; Sgualdo et al., 2015).

98 In this work we present new data including major and trace element analyses on bulk
99 rock, constituent minerals and melt inclusions, as well as Sr-Nd-Pb-He-O isotopes, that are
100 used to constrain: 1) the petrogenesis of picritic melts in terms of physico-chemical
101 conditions and melting regime; 2) the significance of picrite-basalts in the CFB zonal
102 arrangement and 3) the definition of the nature and role of the Afar plume components
103 compared with those of other LIP.

2. Geological setting

The Northern Ethiopian-Yemeni CFB province represents a LIP widely considered to be related to the activity of the Afar plume, on the basis of geophysical, geochemical and volcanotectonic features. (Hoffmann et al., 1997, Ebinger and Sleep, 1998; Pik et al., 1999 and 2006, Davaille et al., 2005; Rogers, 2006; Yirgu et al., 2006; Beccaluva et al., 2009, Bastow et al., 2011; Chang and Van der Lee, 2011; Moucha and Forte, 2011; Natali et al., 2011 and 2013). The impinging of this plume on the Afro-Arabian domain, at least since the Late Eocene, caused up to ca. 3 km regional uplift and doming of the lithosphere (Sengor, 2001; Gani et al., 2007) accompanied by extension and faulting that favoured rapid CFB ascent and eruption. Geophysical data indicate a present day lithosphere thickness of 80-90 km, including 35-40 km of continental crust that is variously underplated by basic intrusions (Bastow et al., 2008 and 2011).

Ethiopian-Yemeni CFB were erupted at ca. 30 Ma (Baker et al., 1996; Hoffmann et al., 1997; Ukstins et al., 2002) as a subcircular (700 km in diameter) originally continuous plateau covering an area of ca. 400,000 km². The basaltic plateau is almost entirely composed by lava flows, (occasionally including dyke swarms; Mège and Korme, 2004), and overlies Mesozoic sedimentary units or the Pan-African crystalline basement (Beccaluva et al., 2009 and references therein). It was subsequently dismembered by the Red Sea/Afar opening (Bellahsen et al., 2003, Bosworth et al., 2005) and is now represented by the uplifted relic portions of the northern Ethiopian and the Yemeni plateaus at the Africa/Arabia conjugate margins (Fig. 1). Restoration of the Yemeni plateau to its original configuration confirms the zonal arrangement of CFB lavas which are represented by 1) low-Ti (LT) tholeiites in the north-western periphery, 2) high-Ti tholeiites (HT1) in the intermediate position and 3) very high-Ti transitional basalts and picrites (HT2) –interbedded with minor HT1 lavas and locally topped by rhyolitic differentiates– near the centre of the subcircular plateau area (Beccaluva

130 et al., 2009; 2011b; Natali et al., 2011). The HT2 lavas, erupted at the inferred Afar plume
131 axis are concentrated in the Lalibela district (Northern Ethiopia) where the volcanic pile is ca.
132 2 km thick and in the Manakhah section at the northernmost portion of the Yemeni plateau,
133 where the maximum thickness is ca. 1 km. GIS-based reconstruction of the HT2 basalt and
134 picrite lavas accounts for ca. 50,000 km³, covering a subcircular area of ca. 200 km diameter,
135 that corresponds to 16% of the whole CFB volume (ca. 310,000 km³). Given the short time
136 span of activity (≤ 1 Ma; Hofmann, 1997) this implies an eruption rate of ca. 0.3 km³ per year
137 for Northern Ethiopia-Yemeni plateau, which is comparable to that of other CFB provinces
138 (Farmer, 2003).

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140 **3. Materials and methods**

141 Sampling has been based on several field works in 2004, 2006, 2009 and 2012 for the
142 Ethiopian plateau and on a field trip organized by the Yemeni Geological Survey in 2007 for
143 the Yemeni counterpart. In particular, HT2 picrite-basalt lavas were mainly collected from
144 the two sections of Manakhah (Northern Yemeni plateau) and Lalibela (Northern Ethiopian
145 plateau). Latitude and longitude of the sampling locations (Table 1) were used for GIS-based
146 reconstructions of the 3D geometry of distinct magmatic groups (LT, HT1 and HT2) carried
147 out using topographic information from NASA SRTM images and applying polynomial
148 interpolations.

149 With regards to the geochemical investigation, the least altered and most representative
150 volcanic lithotypes were selected out of 70 specimens and powdered in an agate mill. Major
151 and trace elements (Ni, Co, Cr, V, Sc, Rb, Sr, Ba and Zr) were analysed by X-ray
152 fluorescence (XRF) on powder pellets, using a wavelength-dispersive automated ARL
153 Advant'X spectrometer at the Department of Earth Sciences, Ferrara University. Accuracy
154 and precision for major elements are estimated as better than 3% for Si, Ti, Fe, Ca and K, and
155 7% for Mg, Al, Mn, Na; for trace elements (above 10 ppm) they are better than 10%. REE, Y,

156 Hf, Nb, Th, and U were analysed by inductively coupled mass spectrometry (ICP-MS) at the
157 Department of Earth Sciences, Ferrara University, using a Thermo-Scientific X-Series.
158 Accuracy and precision, based on the replicated analyses of samples and standards, are
159 estimated as better than 10% for all elements, well above the detection limit.

160 Mineral compositions were obtained at the CNR-IGG Institute of Padova with a Cameca
161 SX-50 electron microprobe (fitted with three wavelength dispersive spectrometers) at an
162 accelerating voltage of 15 kV, and specimen current of 20 nA, using natural silicates and
163 oxides as standards; the analysis of glasses and hydrous mineral phases required a distinct
164 analytical setting and the beam current was lowered to 2 nA in order to mitigate the loss of
165 alkali and halogens. Trace element analyses of constituent mineral phases were carried out at
166 the CNR-IGG of Pavia by laser ablation microanalysis (LAM) ICP-MS, using an Elan DRC-e
167 mass spectrometer coupled with a Q-switched Nd:YAG laser source (Quintel Brilliant). The
168 CaO content was used as an internal standard. Precision and accuracy, better than 10% for
169 concentrations at ppm level, were assessed by repeated analyses of NIST SRM 612 and BCR-
170 2 standards.

171 Sr-Nd-Pb isotopic analyses on bulk rock and handpicked clinopyroxene (~200 mg) were
172 carried out mainly at SUERC (Scottish Universities Environmental Research Centre).
173 Samples were digested with a mixture of HF-HNO₃-HCl and prepared as described by
174 Hardarson et al. (1997). Sr and Nd isotopes were analyzed using a VG Sector 54-30 mass
175 spectrometer in multidynamic mode using exponential corrections for mass fractionation and
176 $^{86}\text{Sr}/^{88}\text{Sr} = 0.1194$, $^{144}\text{Nd}/^{146}\text{Nd} = 0.7219$. Within the period, the NBS 987 standard gave a
177 $^{87}\text{Sr}/^{86}\text{Sr}$ value of 0.710247 (2σ 0.000021, $n=18$) and the internal laboratory J&M standard
178 gave a $^{143}\text{Nd}/^{144}\text{Nd}$ value of 0.511515 (2σ 0.000022, $n=23$). Pb isotopes were analyzed using
179 a Micromass IsoProbe multicollector ICP-MS. Pb was measured using a method similar to
180 that described by Ellam (2006). Mass fractionation was corrected according to the Tl-doping
181 method by adding NIST997 Tl to produce solutions for analysis with 50ng/g Pb and 5 ng/g

182 Pb. The correction was made using an an exponential law and $^{205}\text{Tl}/^{203}\text{Tl}=2.3871$. NBS 981
183 gave $^{206}\text{Pb}/^{204}\text{Pb} = 16.937\pm 4$, $^{207}\text{Pb}/^{204}\text{Pb} = 15.498\pm 5$, and $^{208}\text{Pb}/^{204}\text{Pb} = 36.720\pm 17$ (2σ ,
184 $n=11$). A subordinate subset of Sr-Nd-Pb analyses was carried out at the CNR-IGG of Pisa,
185 where Sr-Nd-Pb isotopic compositions were determined by a Finnigan MAT 262 multi-
186 collector mass-spectrometer following the methods described in Saccani et al. (2013) and Del
187 Moro et al. (2015). In this laboratory replicate analyses of the NBS 987 standard gave an
188 average $^{86}\text{Sr}/^{88}\text{Sr}$ value of 0.710253 ± 13 (2σ , $n=30$), whereas the isotopic standard JNdi-1
189 (Tanaka et al., 2000) gave an average $^{143}\text{Nd}/^{144}\text{Nd}$ value of 0.512098 ± 8 (2σ , $n=25$). As
190 concern Pb, replicate analyses are accurate to within 0.025% (2σ) per mass unit, after
191 applying mass discrimination correction of 0.15 ± 0.01 per mass unit relative to the NBS 981
192 reference composition of Todt et al. (1993).

193 Olivine separates for helium isotope analysis were picked under a binocular microscope
194 and washed ultrasonically in HNO_3 and then water, prior to a final treatment with analar
195 acetone. Helium was extracted from all minerals by in vacuo crushing and analyzed using a
196 MAP 215-50 mass spectrometer at SUERC using procedures slightly modified from Stuart et
197 al. (2000).

198 Oxygen isotopic compositions of mineral separates were measured at the CNR-IGG
199 Institute of Pisa by conventional laser fluorination (Sharp, 1995). To measure the oxygen
200 isotopic composition of each phase, 1–1.5 mg crystals were taken and laser fluorinated with a
201 25W CO_2 laser, at a wavelength of 10.6 μm . Pure fluorine desorbed at 290–310°C from
202 hexafluoropotassium-nickelate salt (Asprey, 1976) was used as reagent. The O_2 produced
203 during laser fluorination was purified from excess fluorine by means of a KCl trap held at
204 180°C, and the chlorine was trapped cryogenically. The gas was then transferred to a 13Å
205 molecular sieve-filled cold finger and analysed for its oxygen isotopic composition using a
206 Finnigan Delta Plus mass spectrometer. Each day of analysis, normally 4–7 aliquots of
207 internal laboratory quartz standards were analysed, with an average reproducibility of ± 0.14

208 % (1σ). Samples were analysed twice, with an analytical precision of ± 0.2 % (1σ) or better.

209 No data correction was necessary for the results, which are reported in the standard $\delta^{18}\text{O}$
210 notation (‰, relative to V-SMOW).

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4. Bulk rock, mineral and melt inclusion chemistry

212 Major and trace element analyses of the investigated HT2 lavas are reported in Table 1,
213 whereas data on their constituent minerals and melt inclusions in olivine, as well as
214 interstitial glasses are reported in Supplementary Table 1 and Table 2, respectively.

215 HT2 basalts and picrites from the Lalibela and Manakhah sequences are always
216 porphyritic in texture with phenocrysts (up to 5 mm across) of olivine, sometimes including
217 Cr-spinel, pinkish clinopyroxene and rare plagioclase; the groundmass is microcrystalline to
218 hypocrySTALLINE (particularly in picrites) and is made up of the same phases with the addition
219 of Fe–Ti oxides, scarce phlogopite (and sporadic kaersutite), alkali feldspar, apatite and
220 interstitial glass. While Ethiopian lavas are generally fresh, Yemeni rocks variably contain
221 secondary minerals mainly represented by serpentine on olivine and infilling carbonates, as
222 indicated by the relatively high Loss On Ignition (LOI, Table 1). The main petrographic
223 difference between picrites and basalts is the prevalence of olivine phenocrysts (up to 20-
224 25%) over clinopyroxene in the former, whereas the reverse is observed in basalts. For
225 classification purposes, picrites can be chemically distinguished from basalts on the basis of
226 their higher MgO (> 13 wt%) and Ni (> 500 ppm) contents. In the Total Alkali-Silica (TAS)
227 diagram of Fig. 2 the HT2 rocks cluster around the alkaline/subalkaline boundary displaying
228 a more alkaline transitional character with respect to LT and HT1 basalts, which mostly plot
229 in the subalkaline field. Within the HT2 group, Yemeni lavas display a more alkaline
230 character with respect to Ethiopian analogues. This cannot be merely ascribed to alteration
231 processes, as the increase of alkalinity in Yemeni lavas is reflected by systematic higher
232 $\text{Na}_2\text{O}/\text{K}_2\text{O}$ ratios (3.5 and 1.9 on average for Yemeni and Ethiopian lavas, respectively) as

234 well as by the clinopyroxene chemistry discussed below. Some relevant intergroup
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2 235 compositional variations are shown in the binary diagrams of Fig. 3 where empirical
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4 236 boundaries discriminate between LT, HT1 and HT2 lavas. The remarkable progressive
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10 238 boundaries, cannot be explained only by variation of degree and P-T conditions of melting
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12 239 but necessarily implies specific enrichment for the respective magma sources. Although less
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14 240 evident, a parallel iron enrichment is observed particularly in HT2 lavas indicating that
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16 241 source variation also involved major element components and were zonally arranged in the
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18 242 Ethiopian-Yemeni plateau (Beccaluva et al., 2009). On the other hand, intra-group variations
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21 243 related to common fractionation processes are shown the by negative correlation between
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23 244 MgO and FeO_t, TiO₂, Nb and Ce. Although occasional crustal contamination can not be
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26 245 excluded as proposed by Pik et al. (1999), particularly for LT magmas, the coherence of the
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29 246 observed trends seems to preclude a major role of this process in the samples considered in
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31 247 this study.

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34 248 Mineral composition in the picrite/basalt suite shows slight but significant variations
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36 249 (Supplementary Table 1a-e). Olivine in picrites and basalts share the same compositional
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39 250 range (Fo from 76.5 to 90.4) with crystal cores of picrites clustering between Fo 87.5 and
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41 251 90.2. NiO in olivines varies from 0.5 to 0.2 wt% and is broadly correlated with the Fo
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44 252 content; CaO is in the range 0.2-0.4 wt% (average 0.3 wt%), comparable with that of olivines
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46 253 in high-MgO lavas from other CFB provinces (Thompson and Gibson, 2000). Clinopyroxene
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49 254 is generally augite in the Lalibela picrites and basalts, whereas it shows a diopside tendency
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51 255 in the Manakhah lavas, reflecting the more alkaline character of the Yemeni rocks.
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53 256 Groundmass plagioclase varies between An₆₂ and An₄₀ and is coupled with minor alkali
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56 257 feldspar (Or up to 53.4), scarce Ti-rich phlogopite (TiO₂ up to 9.5 wt%) and sporadic
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58 258 kaersutite (TiO₂ up to 7.6 wt%). Cr-Fe-Ti oxides are represented by Cr-spinel included in
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61 259 olivine phenocrysts and Ti-magnetite/ilmenite in the groundmass. Cr-spinel (Cr# 0.81-0.65)
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1 260 displays a particularly high TiO₂ content (up to 10.6 wt%) comparable to some other CFB
2 261 picritic rocks (Desta et al., 2014 and references therein). Groundmass Ti-magnetite shows
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4 262 ülvospinel content varying from 34 to 56% and is always coupled with ilmenite.
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7 263 Melt inclusions (MI, 20-100 µm in size) in olivine phenocrysts were analyzed in two
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9 264 picrites (LAL70, LAL73) and one basalt (LAL42) from Lalibela (Table 2). Although
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11 showing a wide compositional range (MgO 3.3-15.4 wt%) they invariably display high TiO₂
12 265 content (2.3-5.4 wt%). In the TAS diagram of Fig. 2 MI vary in composition from picrite to
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14 266 basaltic andesite, mostly overlapping the picrite-basalt bulk rock distribution. The analysis of
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16 267 MI-LAL70, corrected for post-entrapment olivine crystallization (Danyushevsky et al., 2002),
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18 268 shows a picritic composition (MgO = 15.4 wt%) and displays a slightly more alkaline
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20 269 character compared to HT2 Ethiopian bulk rocks. Analyses of MI-LAL73 and MI-LAL42
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22 270 show a variable composition in the basaltic field, which reflects trapping of cogenetic melt
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24 271 fractions at various differentiation stages. It is noteworthy that groundmass interstitial glasses
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26 272 from picrite LAL73 (Table 2) expand the MI fractionation trend toward an andesitic
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28 273 composition, thus recording a wide magmatic evolution even in the same sample.
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36 275 The trace element characteristics of Northern Ethiopia-Yemeni HT2 picrite-basalt lavas
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38 276 are illustrated in the Chondrite (Ch)-normalized and Primordial Mantle (PM)-normalized
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40 277 diagrams of Fig. 4. The Rare Earth Elements (REE) distribution invariably shows a
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42 278 continuous positively fractionated pattern with remarkable Light (L)REE enrichment
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44 279 (La_N/Yb_N from 8 to 18; Tb_N/Yb_N 2.7-3.8). Although subparallel and largely overlapped, the
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46 280 lowest and the highest REE levels are recorded by some picrites and basalts, respectively
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48 281 (Fig. 4a). In Fig. 4b the PM-normalized incompatible element distribution of HT2 lavas
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50 282 generally shows upward-convex patterns that conform to those of transitional/alkaline Ocean
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52 283 Island Basalt (OIB) from within-plate setting (Sun and McDonough, 1989; Hoffmann, 1997),
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54 284 and, particularly to the EM2-types (Weaver et al., 1991). However, these spiderdiagrams
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2 285 reveal that most HT2 rocks are enriched in Ti and more in general in High Field Strength
3 286 Elements (HFSE) such as Zr and Nb.

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5 287 The incompatible element distribution of MI satisfactorily matches the picrite-basalt bulk-
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7 288 rock field, with sample MI-LAL70 approaching the most primary picritic compositions,
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9 289 whereas the other MI and the interstitial glass depict variably enriched patterns which reflect
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11 290 the fractionation degree (Fig 4c). Trace element analyses of clinopyroxene in the picrite-
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13 291 basalt association are reported in Table 3 and in the Ch-normalized diagram of Fig. 5. The
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15 292 observed upward-convex patterns (La_N/Yb_N 1.0-4.4, Sm_N/Yb_N 3.3-10.6) resemble those from
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17 293 alkaline basalts (Jeffries et al., 1995), in agreement with the transitional to alkaline character
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19 294 of the relative magmas, which is even more emphasized in the Yemeni rocks. The HREE
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21 295 fractionation (Tb_N/Yb_N 2.1-6.0) of these clinopyroxenes mimics those of the relative bulk
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23 296 rocks. Similar clinopyroxene patterns have been also observed in High-MgO plume-related
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25 297 magmas from other LIPs (Kamentesky et al. 2012; Wei et al., 2015) and interpreted as a
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27 298 garnet signature in the magma sources. Interestingly, significant clinopyroxene compositional
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29 299 variations (e.g. height of the pattern) are often observed within the same rock sample,
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31 300 possibly as a result of either polybaric fractionation or mixing/cumulus processes.
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41 302 **5. Petrogenesis of the picrite-basalt association**

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43 303 The identification of primary and parental magmas of a suite is not straightforward even
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45 304 when dealing with high-MgO lavas, not only because of the difficulty of estimating
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47 305 thermobarometric conditions, melting regime and source composition, but also due to mixing
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49 306 and cumulus processes affecting magmas rising in a short time along the same feeding
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51 307 system, as is the case of the investigated picrite-basalt association. The integration of multiple
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53 308 approaches, such as major element phase equilibria and trace element modelling, may help to
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55 309 put constraints on primitiveness of magmas and P-T-X conditions of their mantle sources as
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57 310 discussed below.
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2 312 *5.1 The quest for primary and parental melts*
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5 313 Most picrites of the studied HT2 association have mg# between 0.68 and 0.72, thus
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7 314 showing potential equilibrium with a mantle source with olivine Fo>0.88 (Herzberg and
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9 315 O'Hara, 2002, Green and Falloon, 2005). Stronger constraints are provided by the MgO-FeO
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11 316 diagram (Fig. 6, modified after Herzberg et al., 2007), where the petrogenetic grid for
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13 317 primary melts and fractionation trends of OIB, plateau basalts and MORB suites are reported,
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15 318 together with Fo composition arrays of coexisting olivine. In this diagram, the Ethiopian-
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17 319 Yemeni HT2 lavas define a relatively high FeO distribution roughly reflecting various extent
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19 320 of magmatic fractionation from picrites to basalt, as well as significant olivine accumulation
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21 321 in some samples (samples labelled with *c*). Coherently, in Fig. 7, it can be observed that
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23 322 although many samples show an olivine compositional range encompassing the equilibrium
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25 323 conditions with the respective whole rock ($K_{d_{\text{Ol-rock}}} = 0.28-0.32$), most of the analyzed
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27 324 olivines show disequilibrium. This is probably due to the intrinsic mechanism of olivine
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29 325 crystallization in picritic magmas, which is characterized by rapid growth and compositional
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31 326 changes, leading to persistent disequilibrium conditions during crystallization. This implies
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33 327 that only a few picritic samples (*i.e.* LAL5, LAL6 and LAL9), displaying restricted olivine
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35 328 compositional ranges in equilibrium ($\text{Fo}_{88.8-90.2}$) with the respective bulk rock (mg# 0.69–
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37 329 0.72), confidently represent suitable parental magmas for petrogenetic modelling
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39 330 Accordingly, the accumulated fractional melting model of Herzberg and Asimow (2015) was
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41 331 applied to these most primitive picrites, assuming $\text{Fe}^{2+}/\text{Fe}_{\text{tot}} = 0.9$, using as source
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43 332 composition both the notional garnet lherzolite source KR4003 (Walter, 1998) and the
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45 333 average of Ethiopian lherzolite mantle xenoliths (MgO 39.7 wt%, FeO 8.6 wt%; Beccaluva et
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47 334 al., 2011b). In all cases no limit to the model applicability was observed for the selected
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49 335 compositions (see Appendix A in Herzberg and Asimow, 2015): the calculated primary melts
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51 336 have MgO contents in the range 19.8-21.6 wt% and average mantle potential temperatures
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337 nearly coincident either assuming as magma source the Ethiopian xenoliths (T_p 1570°C) or
338 the KR4003 peridotite (T_p 1585 °C). This suggests that, at least in terms of bulk chemical
339 composition, the sub-Ethiopian mantle xenoliths are appropriate for modelling the source of
340 the Ethiopian-Yemeni CFB.

341 Compared with other high-MgO lavas from various plateau basalts, OIB and MORB
342 settings (Fig. 8), the Ethiopian-Yemeni data overlap those obtained from NAP and Deccan
343 and, together with other CFB provinces (Etendeka, Karoo, Emeishan and Siberia), represent
344 the highest T_p range (1550-1700°C) which characterizes Phanerozoic continental LIP. By
345 contrast, oceanic plateaux, such as Ontong Java and Caribbean, and OIB (Hawaii, Gorgona
346 and Iceland), are generally cooler in a T_p range 1550-1400°C. Application of the above
347 modelling to the most primitive LT and HT1 basalts allows comparison between the
348 peripheral magmas of the Ethiopian-Yemeni plateau with those of its axial zone (HT2 group):
349 the calculated LT and HT1 primary magmas (data from Baker et al., 1996; Beccaluva et al.,
350 2009; Teklay et al., 2005 and unpublished author data) have 14 and 18 wt% MgO,
351 respectively (Fig. 6). The corresponding T_p are 1430°C for LT, and 1530°C for HT1 which
352 indicates that the Ethiopian-Yemeni CFB are clearly zoned both compositionally and
353 thermally (ΔT_{pmax} 140°C). Overall, in spite of the limits of the applied model for the T_p
354 estimate (see discussion by Heinonen et al., 2015), the obtained data are reasonably
355 consistent both for the internal comparison within Ethiopian-Yemeni CFB and with other LIP
356 occurrences worldwide. Noteworthy, alternative thermometric estimates (such as Al-in-
357 olivine thermometer, Coogan et al., 2014), although giving systematically lower temperatures
358 (ca. 100°C), do not obscure the presence and extent of LIP-related thermal anomalies
359 (Heinonen et al., 2015).

360 Phase equilibria constraints, according to Herzberg et al. (2007), indicate that the
361 calculated primary HT2 picrites correspond to a melt fraction in the range 0.10-0.15
362 generated by a garnet-bearing peridotite mantle source at 4-5 GPa pressure. The application

363 of the Gudfinnsson and Presnall (2005) and Herzberg and Zhang (1996) whole rock
364 geobarometers confirm the above pressure range of melting. On the other hand, the same
365 parameterization indicates that LT and HT1 primary magmas could be generated by 0.15-
366 0.25 melt fraction of a spinel peridotite in the pressure range < 2 to 3 GPa, which extends the
367 previous estimates for LT and HT1 generation to higher pressure (1.3-2.2 GPa; Beccaluva et
368 al., 2009).

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370 *5.2 Modelling the fractionation trends*

371 This section will discuss the application of a forward fractionation model to test the
372 possible derivation of HT2 basalts from the associated picrites. Starting from the calculated
373 primary picrites, fractional crystallization modelling (*Petrolog v.3*; Danyushevsky and
374 Plechov, 2011) indicates that the observed differentiation trends up to the most fractionated
375 basalt (MgO ca. 5 wt%) can be accounted for by relatively low pressure fractionation (Fig. 6
376 and 8). In fact, high pressure conditions would not fit mineral/rock equilibria, whereas
377 polybaric fractional crystallization starting at pressure ≤ 1 GPa can reproduce the
378 compositional variations observed in olivine and particularly clinopyroxene phenocrysts, as
379 well as in the whole rocks. Accordingly, equilibrium olivine starts to crystallize at a
380 temperature of ca. 1510°C (Fo 91.3), implying that primary magmas (T_p 1570°C) upwelled
381 nearly adiabatically to the base of the crust. Crystal fractionation is dominated by olivine up
382 to ca. 1260°C where clinopyroxene (mg# ca. 0.86) join olivine in cotectic crystallization (Fig.
383 8). The final basaltic composition is attained after removal of a total amount of ca. 50% solid
384 phases (33 % olivine and 17% clinopyroxene).

385 The above data suggest that the genesis and evolution of the picrite-basalt association
386 reflect a complex T-P-X path, likely starting from the generation of primary picritic melts at
387 high potential temperature and pressure (1570°C, 4-5 GPa), followed by nearly-adiabatic
388 magma rising up to the onset of non-adiabatic regime at crustal level (<1 GPa), where picritic

389 magmas slowed down and underwent fractional crystallization with formation of variably
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2 390 differentiated magmas. The same modelling has been also applied for LT and HT1 primary
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4 391 melts (data from Baker et al., 1996; Beccaluva et al., 2009; Teklay et al., 2005 and
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7 392 unpublished author data). For LT olivine (Fo 90.3) starts to crystallize at 1420°C, followed
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9 393 by plagioclase (An₇₇₋₈₁) at 1200°C, and clinopyroxene (mg# 81-82) at ca. 1160°C. Removal
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11 394 of ca. 60% solid phases (Ol 26%, Pl 26%, Cpx 8%) is required to reach the composition of
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13 395 the most fractionated LT basalt (Fig. 8). Fractionation of HT1 magmas displays intermediate
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15 396 features between LT and HT2 (Fig. 6), with the most fractionated basalt resulting from
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17 397 removal of ca. 50% solid phases (Ol 29%, Cpx 17%, Pl 6%). It is noteworthy that, while the
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19 398 HT2 picrite-basalt evolutionary trend is characterized by the appearance of clinopyroxene
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21 399 crystallization after olivine and a mild Fe-enrichment, conforming to a transitional to alkaline
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23 400 affinity, the LT trend is typically tholeiitic and characterized by extensive plagioclase
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25 401 crystallization after olivine, leading to a marked Fe-enrichment of the evolved magmas (inset
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27 402 of Fig. 6).
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34 403 Trace element fractional crystallization modelling has also been carried out on the basis
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36 404 of the amount of fractionated solid phases calculated from major elements. Model results
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38 405 reported in PM-normalized incompatible element diagrams of Fig. 9, indicate that there is a
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40 406 good agreement between the computed and the observed differentiated basalts using
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42 407 appropriate partition coefficients (K_d) from GERM database for Ol, Cpx and Pl. It is
43
44 408 important to note that nearly identical results are obtained using the clinopyroxene K_d
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46 409 calculated as cpx phenocryst/bulk rock ratios (Table 1 and 3).
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51 410 52 53 411 *5.3 Trace element modelling for the generation of HT2 primary melts* 54

55 412 The trace element distributions of the Ethiopia-Yemeni HT2 picrite-basalt are compared
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57 413 in Fig. 10a with other high-Mg lavas from plateau basalts and OIB. Clearly, the differences in
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59 414 incompatible element enrichment among the various provinces are too wide to be accounted
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415 for by olivine cumulus/fractionation processes which commonly affect picritic magmas.
416 Therefore, these differences have to be ascribed to variations in source compositions and/or
417 melting modes. As indicated in the previous section and already suggested by some authors
418 (Kamenetsky et al., 2012; Heinonen & Luttinen, 2008; Gibson et al., 2000), the marked
419 LREE enrichment, the HREE fractionation and the high V/Sc ratios in MgO-rich lavas from
420 most CFB provinces conform to magma originating from deep and variably metasomatized
421 mantle sources with residual garnet. On the other hand, high-Mg lavas from the Caribbean
422 plateau and NAP, together with those from Iceland and Gorgona do not show incompatible
423 element enrichment and HREE fractionation, thus suggesting shallower melting (spinel
424 facies) of their mantle sources (Rèvillon et al., 1999 and 2000).

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425 However, the above considerations are not sufficient to justify the Ti-(Fe) and HFSE (e.g.
426 Zr, Hf) enrichments recorded in CFB lavas such as those from Ethiopia-Yemen, Siberia,
427 Karoo and Emeishan, thus requiring additional components in their mantle sources. The
428 nature of these components could be attributed either to primordial planetary differentiation
429 (Elkins-Tanton et al., 2002; 2011; Solomatov, 2016 and references therein), to plume-related
430 pyroxenite cumulates (Rogers et al., 2010), or to pyroxenite/eclogite material deriving from
431 ancient subducted oceanic slabs subsequently incorporated and melted at high pressure in a
432 rising mantle plume (Siberia, Sobolev et al., 2007 and 2009; Northern Ethiopia, Beccaluva et
433 al., 2009; Karoo, Heinonen et al., 2008). This last hypothesis has gained ground, particularly
434 for explaining the CFB magma genesis. Accordingly, Jennings et al. (2014a; 2014b) observed
435 that the peculiar features of CFB ferropicrites require high $P-T_p$ melting beneath the
436 continental lithosphere that preferentially taps the most fusible eclogite/pyroxenite-rich
437 sublithospheric mantle lithologies.

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438 In this context we extensively reviewed the pyroxenite/eclogite compositions from the
439 literature that should be added to the sub Ethiopian mantle (assumed as lherzolite xenolith
440 GOJ40A; Beccaluva et al., 2011b) in order to obtain a suitable source for the HT2 primary

441 magmas. Best fit model calculations (Fig. 10b) show that primary picrites could be generated
442 by ca. 10 % of both batch and fractional melting of a mixed garnet lherzolite (85%) and
443 eclogite (15%) source, with the latter corresponding to Neoproterozoic eclogite from Hoggar
444 (Berger et al., 2014). Noteworthy, these eclogites, chemically resembling transitional MORB
445 (cf Le Roex, 1987), record subduction episodes of ancient oceanic lithosphere involved in the
446 Pan-African tectonomagmatic events all along the North African-Arabian domain (Stern and
447 Johnson, 2010). Conceptually, these results can be accommodated in a multi-stage
448 petrogenetic process in which the source is a piclogite domain (Bianchini et al., 2010),
449 resulting from the interaction of pyroxenite/eclogite melt with a garnet peridotite matrix. This
450 implies pyroxenite mobilization, melt generation, upward migration and crystallization,
451 during the incipient rising of the plume head, which is followed by shallower melting of
452 piclogite domains due to further upwelling, with generation of HT2 magmas at 4-5 GPa
453 pressure (cf Tuff et al., 2005; Sobolev et al., 2009; Rosenthal et al. 2014). Although
454 simplified for what concern the physical-chemical aspects of the eclogite component
455 mobilization, the trace element modelling allows placing robust constraints on the extent of
456 the peridotite-eclogite hybridization, the degree of melting and the presence of residual garnet
457 in the HT2 mantle source. Moreover, these results are in good agreement with those from
458 major element modelling, as well as with the isotopic data that will be discussed in the next
459 section. Further investigations and modelling on the eclogite/peridotite interaction are beyond
460 the scope of this work and will be the object of future researches.

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6. Sr-Nd-Pb-He-O isotope systematics

463 Sr-Nd-Pb isotopic data of bulk rock and clinopyroxene of the Ethiopian-Yemeni CFB
464 (LT, HT1, HT2) are reported in Table 4. As a whole, the Ethiopian-Yemeni CFB show the
465 following ranges: $^{87}\text{Sr}/^{86}\text{Sr}$ 0.7032-0.7044 and $^{143}\text{Nd}/^{144}\text{Nd}$ 0.51274-0.51306, and $^{206}\text{Pb}/^{204}\text{Pb}$
466 18.02-19.41, $^{207}\text{Pb}/^{204}\text{Pb}$ 15.48-15.63, $^{208}\text{Pb}/^{204}\text{Pb}$ 37.55-39.46, with the Yemeni rocks having

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2 468 the higher lead isotopic compositions. Sr-Nd-Pb isotopic data are reported in Fig. 11 and
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5 469 compared with the composition of other flood basalts from major Phanerozoic LIPs. In Fig.
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7 470 11a the Ethiopian-Yemeni CFB are extremely clustered with $\epsilon_{\text{Sr}}^{(t)}$ between and -18.0 and -2.1,
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9 471 and $\epsilon_{\text{Nd}}^{(t)}$ between +8.3 and +2.3, showing partial overlap with the upper-left portion of CFB
10 472 fields from Karoo, Deccan and Siberia with the Parana-Etendeka CFB field slightly displaced
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12 473 in the opposite direction. Therefore, very radiogenic Sr and unradiogenic Nd isotopic
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14 474 compositions are lacking in the Ethiopian-Yemeni lavas, implying the absence of the
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16 475 “continental crust” fingerprint envisaged in other CFB provinces. Further comparison is
17 476 provided by Fig. 11b, where the $^{207}\text{Pb}/^{204}\text{Pb}$ vs $^{206}\text{Pb}/^{204}\text{Pb}$ composition of Ethiopian-Yemeni
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19 477 lavas is plotted together with data from other Phanerozoic oceanic and continental LIP
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21 478 worldwide (Jackson and Carlson, 2011 and references therein). Here, LT lavas largely
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23 479 overlap those of Baffin Island-West Greenland, Ontong Java Plateau, Deccan and Siberia,
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25 480 whereas HT2 lavas are significantly displaced on the right of the geochrons. According to
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27 481 Jackson and Carlson (2011), this compositional shift could be related to the incorporation of
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29 482 recycled oceanic crust into the non-chondritic Primitive Mantle reservoir, in agreement with
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31 483 the trace element modelling discussed in the previous section.
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39 483 Differences within the Ethiopia-Yemen CFB are emphasized in Fig. 12 which also
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41 484 includes the notional OIB mantle end-members (DM, EM1, EM2, HIMU; Hofmann, 1997;
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43 485 Stracke et al., 2005). There is a clear covariance between Pb and Sr isotopes, with HT2 lavas
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45 486 showing the more radiogenic values ($^{87}\text{Sr}/^{86}\text{Sr}$ 0.7040-0.7045, $^{206}\text{Pb}/^{204}\text{Pb}$ 18.8-19.4,
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47 487 $^{208}\text{Pb}/^{204}\text{Pb}$ 38.6-39.5), while the Yemeni rocks plot out of the trend due to more radiogenic
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49 488 lead. A covariance is also observed for TiO_2 vs $^{87}\text{Sr}/^{86}\text{Sr}$ and $^{206}\text{Pb}/^{204}\text{Pb}$, (Fig. 13) with a clear
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51 489 distinction between LT, HT1 and HT2 lava groups. As already seen for the incompatible
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53 490 element distribution, this indicates geochemically distinct mantle sources for LT, HT1 and
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55 491 HT2 lavas, in relation to a progressive enrichment of the titaniferous geochemical component
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57 492 which characterizes the Afar plume. The above correlation (Fig. 13), together with the
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2 494 relatively restricted range of isotopic data taken as a whole (Fig.11) further precludes the
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5 495 occurrence of significant crustal contamination processes for the samples considered in this
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7 496 study.

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10 497 For the Yemeni HT2 lavas, the observed differences in lead isotopic composition may be
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12 498 inherited from ancient components in the lithosphere (HIMU?) partially masked (but not
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14 499 reset) by the Afar plume metasomatism. It is noteworthy that these new Sr-Nd-Pb isotopic
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16 500 compositions encompass those of Yemeni flood basalts considered by Baker et al., (1996 and
17 501 2000) unaffected by crustal contamination.

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19 502 Additional information can be obtained comparing the Sr-Nd-Pb isotopic compositions of
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21 503 bulk rock and constituent clinopyroxene, available for a subset of HT2 lavas including both
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23 504 basalts and picrites. The isotopic misfit variously observed in the clinopyroxene-bulk rock
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25 505 pairs, cannot be simply ascribed to differential radiogenic ingrowth, as demonstrated by the
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27 506 $^{206}\text{Pb}/^{204}\text{Pb}_{(i)}$ vs $^{87}\text{Sr}/^{86}\text{Sr}_{(i)}$ diagram reporting the initial isotopic values calculated at 30 Ma
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29 507 (Fig. 12d). The most extreme differences between bulk rock and coexisting clinopyroxenes
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31 508 are observed particularly in those picrites that clearly reflect crystal (ol+cpx) accumulation
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33 509 such as LAL70 and LAL73. For these samples, clinopyroxene crystals should not be
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35 510 considered real phenocrysts and could be interpreted as “antecrysts”, i.e. phases that although
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37 511 grown in the same magmatic system did not crystallize directly from the host magma, thus
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39 512 recording marked isotopic disequilibrium. Similar disequilibrium evidences between cpx
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41 513 crystals and host rocks have been observed in other plume related magmas (Wei et al., 2015
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50 and references therein).

51 514 Olivine phenocrysts from HT2 lavas (Table 5) have $^3\text{He}/^4\text{He}$ that range between 9.6 and
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53 515 19.1 R_A with more than 50% of the analysed samples in the range 15.0-19.1 R_A . These values
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55 516 are significantly higher than those commonly observed in MORB ($8 \pm 1 R_A$, Graham, 2002),
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57 517 in Ethiopian LT basalts ($< 5 R_A$) and in lavas from other African volcanic districts (Marty et
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59 518 al., 1996; Pik et al., 2006). Moreover, they are also higher than those recorded in mantle
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2 520 xenoliths from the Afro-Arabian system (mostly in the range 6.5-8.0 R_A) which represent the
3 local lithospheric mantle (Aulbach et al., 2011; Beccaluva et al., 2011b; Sgualdo et al., 2015).
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5 521 This He signature can be attributed to a deep, relatively undegassed mantle, although it is
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7 522 lower than the ratios measured in the Proto-Iceland plume picrites (up to 50 R_A , Stuart et al.,
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9 523 2003; Starkey et al., 2009). This may be the result of mixing with a low $^3\text{He}/^4\text{He}$ mantle end-
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11 524 member, also including recycled eclogite ($< 0.1 R_A$, Day et al., 2015).

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14 525 Oxygen isotopic composition of olivine and clinopyroxene phenocrysts (Table 6) displays
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16 526 $\delta^{18}\text{O}$ values between 5.2 and 6.9 ‰ that are higher than those generally accepted for mineral
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18 527 phases in mantle derived magmas (Mattey et al., 1994). Similar values have been recorded in
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20 528 Yemeni CFB by Baker et al. (2000) and interpreted as the result of shallow level crustal
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22 529 contamination. However, in our opinion, they could be better attributed to the involvement of
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24 530 pyroxenite/eclogite lithologies in magma sources, as modelled in the previous section and
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26 531 suggested by $\delta^{18}\text{O}$ data on eclogite and metasomatized mantle xenoliths (Deines and
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28 532 Haggerty, 2000; Hao et al., 2015).

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32 533 The possible involvement of Pan-African eclogite incorporated in the Afar plume has
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34 534 been tested using the available Sr and Nd isotopic data included in Berger et al. (2014) and
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36 535 those from lherzolite xenoliths from sub Ethiopian mantle (Beccaluva et al., 2011b). Mass
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38 536 balance calculation between these two end-members (lherzolite $^{87}\text{Sr}/^{86}\text{Sr} = 0.7029$,
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40 537 $^{143}\text{Nd}/^{144}\text{Nd} = 0.5132$; eclogite $^{87}\text{Sr}/^{86}\text{Sr} = 0.7056$, $^{143}\text{Nd}/^{144}\text{Nd} = 0.5128$) indicates that a
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42 538 mixture of ca. 11-13% eclogite and 89-87% lherzolite satisfactorily match the Sr-Nd isotopic
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44 539 composition of HT2 lavas. Unfortunately, Pb isotopes are not available for these eclogites but
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46 540 radiogenic Pb isotopic composition, compatible with that inferred for HT2 magma sources,
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48 541 have been reported for other eclogitized garnet-pyroxenites included in Pan-African terranes
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50 542 (Schmädicke et al., 2015). These mixture proportions are also compatible for the oxygen
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52 543 isotopes if we assume 5.2 as $\delta^{18}\text{O}$ value for lherzolite (Mattey et al., 1994) and 9.0 for
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54 544 pyroxenite/eclogite recycled component (Pearson et al., 1991). We may conclude that, based
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1 545 on the available data, the source mixture resulting from isotopic mass-balance is in
2 546 reasonable agreement with that inferred from trace element modelling discussed in the
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4 547 previous section (15% eclogite, 85% lherzolite).
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9 549 **7. Discussion and Conclusions**

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12 550 Detailed petrological and geochemical studies on the HT2 picrite-basalt lavas erupted at
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14 551 the core of the Ethiopian-Yemeni CFB plateau, and comparison with the associated HT1 and
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16 552 LT basalts, allow the thermo-compositional effects induced by the Oligocene Afar plume in
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18 553 the respective magma sources to be constrained.
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22 554 Major element modelling and phase equilibria constraints on selected HT2 lavas indicate
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24 555 that picritic primary melts formed by partial melting of a garnet bearing mantle source in the
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26 556 pressure range 4-5 GPa, at potential temperature (T_p) of 1570°C. Fractional crystallization
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28 557 modelling (Fig. 6 and 8) shows that the observed differentiation trends from primary picrites
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30 558 to the most fractionated basalts can be accounted for by relatively low pressure polybaric
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32 559 fractionation, which likely occurred from the base of the continental crust (< 1 GPa), where
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34 560 upward advection of magma slowed down allowing more effective cooling. Accordingly,
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36 561 equilibrium olivine should appear on the liquidus at temperature of ca. 1510°C (Fo 91.3), as a
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38 562 result of nearly adiabatic uprising of primary magmas from the source region (T_p 1570°C,
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40 563 P_{\min} 4 GPa, 0.6°C/km adiabatic gradient) up to the Moho (35-40 km). The most differentiated
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42 564 basaltic compositions were satisfactorily reproduced after removal of ca. 50% solid phases
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44 565 represented by olivine and clinopyroxene, the latter appearing on the cotectic at ca. 1260°C,
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46 566 *i.e.* at the picrite-basalt transition. However, it is noteworthy that the compositional variability
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48 567 of the HT2 lavas (and MI in olivine) appears to be related to widespread
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50 568 accumulation/disequilibrium crystallization and mutual magma mixing at crustal levels. This
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52 569 is also indicated by the wide $Kd_{Ol/Cpx-melt}$ range, the cross-cutting incompatible element cpx
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54 570 patterns, as well as the significant isotopic differences between Cpx and the host bulk rock
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2 572 often recorded in HT2 lavas. This scenario conforms to a high magma production rate of
3 basic melts rising along the same feeding system at the plume axial zone.

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5 573 The thermal regime of the plateau outer zones was significantly cooler with respect to the
6 inner part, as indicated by P-T genetic conditions of LT primary magmas (T_p 1430°C, <2 to 3
7 574 GPa), thus suggesting a thermal difference of 140°C between the axial and peripheral zones
8 of the plume head. The spatial distribution of the Ethiopian-Yemeni CFB, and their P-T-X
9 575 genetic conditions allow to reconstruct the Oligocene configuration of the Afar plume head as
10 a melting region where plume-related thermo-chemical effects decreased outward (Fig. 143).
11 In particular, the anomalously high Ti, together with Fe and incompatible element
12 576 enrichments, of the HT2 picrite-basalt suite imply the presence of peculiar components that
13 enriched the relative mantle sources. A careful review of Fe-Ti reservoirs -such as
14 577 pyroxenites/eclogites included in the upper mantle lithologies and/or relics of subducted
15 oceanic lithosphere (*e.g.* ophiolites)- reveals that the best candidate to hybridize the HT2
16 578 mantle source is represented by eclogites derived from Transitional-MORB protoliths
17 subducted during the Neo-Proterozoic, and currently outcropping in Pan-African terranes
18 (Berger et al., 2014). Model calculations indicate that HT2 primary magmas could be
19 579 generated by ca. 10% partial melting of a “piclogite” (*cf* Bianchini et al., 2010 and references
20 therein) mantle source composed by 85% garnet lherzolite and 15% of the above mentioned
21 580 eclogite. The inferred source characteristics do not conflict with the subchondritic $^{187}\text{Os}/^{188}\text{Os}$
22 ratio observed in HT2 Ethiopian picrites by Rogers et al. (2010), who limit to 15% and $\leq 1\text{Ga}$
23 in age, the possible contribution of Neo-Proterozoic subduction-related material in the source
24 581 of the Ethiopian picrites. Accordingly, recent seismic tomography investigations seem to
25 confirm the presence of a low velocity zone, at ca. 500 km beneath the Afar (Thompson et
26 582 al., 2015), that in our opinion could represent piclogite domains stored at the mantle
27 transition zone, partially remobilized and entrained by the Oligocene plume, and still
28 583 persisting beneath the region.

597 The isotopic analyses extensively carried out on LT, HT1 and HT2 lavas reveal that while
1
2 598 $^{143}\text{Nd}/^{144}\text{Nd}$ is rather homogeneous in the various magma types, Sr-Pb isotope ratios are
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4
5 599 positively correlated with TiO_2 reflecting the relative contribution of the pre-existing
6
7 600 lithosphere and incoming sub-lithospheric plume components which increase from peripheral
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9
10 601 LT to the axial zone HT2 lavas (Fig. 13). Accordingly, the Afar plume fingerprint is
11
12 602 particularly recorded by HT2 picrite-basalt magmas that are characterized by a limited
13
14 603 isotopic variability, and cluster between the notional DM and EM2 mantle end-members.
15
16
17 604 Coherently, this isotopic signature is recorded in mantle xenoliths from the plateau area
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19 605 (Beccaluva et al., 2011b; Bianchini et al., 2014) and still persists in the recent Afar volcanism
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21
22 606 (Rooney et al., 2012). By contrast, Neogene volcanics and associated mantle xenoliths from
23
24 607 Afro-Arabian domains located outside from the Ethiopian-Yemeni plateau display a much
25
26
27 608 wider isotopic spectrum mainly varying between the DM, HIMU and EM1 end-members
28
29 609 (Furman et al., 2006; Rogers, 2000; Beccaluva et al., 2011b; Bianchini et al., 2014; Sgualdo
30
31 610 et al., 2015).

32
33
34 611 The new He isotopic data are in accordance with literature data that recorded $^3\text{He}/^4\text{He}$ up
35
36 612 to 20 R_A in the same lava type (Marty et al., 1996; Pik et al., 2006). We tentatively explain
37
38
39 613 this as a mixing of mantle plume, “normal” mantle domains and recycled eclogite. The
40
41 614 involvement of the latter is corroborated by the recorded O isotopic compositions that display
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44 615 values that are higher than those generally observed in magmas derived from purely
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46 616 peridotitic mantle.

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49 617 Therefore, the Ethiopian-Yemeni province may represent a typical well constrained
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51 618 example of a LIP generated by a hot mantle plume. It entrained eclogitic and undegassed
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54 619 components, ultimately leading to a plume head with decreasing thermo-chemical effects
55
56 620 over a radial extension of ca. 400 km, from which zonally arranged CFB were generated.

57
58 621 The worldwide comparison of high-MgO lavas from various CFB Phanerozoic provinces
59
60
61 622 (Fig. 8 and 10a) indicates that these magmas shares high mantle potential temperatures (T_p)
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1550-1700°C) supporting the existence of plume related thermal anomalies as triggering events for the rapid emplacement of large volume of basic magmas, that often preceded continental break-up and oceanic opening. However, thermal anomalies are not necessarily accompanied by strong geochemical signature (e.g. Ti and incompatible element enrichments), thus suggesting that mantle plumes exert variable, sometimes weak, chemical effect on the lithospheric CFB magma sources. This could explain the difficulty to recognize a precise plume fingerprint in some CFB occurrences as is the case of ferropicrite dykes from Vestfjella, that are at least in part isotopically indistinguishable from MORB of the SW Indian Ridge (Heinonen et al., 2010). Similarly, low-Ti and high-Ti CFB of the Parana province (ca. 130 Ma) although overlapping the “Tristan plume” isotopic signature (Thompson et al., 2001), also record significant compositional and isotopic (Sr-Nd) similarities with Late Archean-Proterozoic and Early Jurassic (CAMP, ca. 200 Ma) tholeiites from the South American platform (Iacumin et al., 2003). We may conclude that, although the mantle plume hypothesis represents a suitable explanation for generating many LIP characterized by high-T and high eruption rate of CFB, the reliable assignment of specific geochemical components to mantle plume effects should be based on extensive and detailed investigations of the spatial-temporal distribution and thermo-compositional features of the entire magmatic province.

641

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36 **987 Table captions**

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41 **989 Table 1** – Bulk rock major and trace element composition of HT2 basalts and picrites from
42 the Ethiopian-Yemeni CFB.
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45 **991 Footnote:** $\text{Fe}_2\text{O}_3/\text{FeO} = 0.15$; $\text{mg\#} = (\text{MgO})/(\text{MgO}+\text{FeO})$ mol%. * = analyses from Beccaluva
46 et al. (2009) and Natali et al. (2011). ** = corrected for typing error in Beccaluva et al.
47 (2009).
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55 **995 Table 2** – Major (EMPA) and trace element (LA-ICP-MS) averaged analyses of melt
56 inclusions (MI) in olivine and interstitial glass (GI) from Ethiopian HT2 basalts and picrites.
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998 **Table 3** – LA-ICP-MS trace element analyses of clinopyroxene in HT2 picrites and basalts
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2 999 from Ethiopian-Yemeni CFB.
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7 1001 **Table 4** – Sr-Nd-Pb isotopic composition of LT, HT1 and HT2 lavas and separated
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9 1002 clinopyroxenes from the Ethiopian-Yemeni CFB.
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12 1003 **Footnote:** Sr, Nd and Pb initial isotope ratios, calculated at 30 Ma, are also reported.
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17 1005 **Table 5** – $^3\text{He}/^4\text{He}$ composition of olivine from Ethiopian HT2 lavas.
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19 1006 **Footnote:** Measured ratios (R) are normalized to the atmospheric ratio (Ra; 1.384×10^{-6}).
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22 1007 Uncertainties are expressed as 1σ .
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27 1009 **Table 6** – $\delta^{18}\text{O}$ of olivine and clinopyroxene separated from Ethiopian HT2 lavas.
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29 1010 **Footnote:** Uncertainties are expressed as 1σ .
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34 1012 **Supplementary Table 1** – Representative analyses of constituent minerals from HT2 picrite-
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36 1013 basalt lavas. a) olivine, b) clinopyroxene, c) feldspars, d) Fe-Ti-Cr oxides, e) phlogopite and
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38 1014 amphibole.
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49 1018 **Figure captions**
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54 1020 **Figure 1** – Sketch map of the Oligocene Ethiopian-Yemeni CFBs (modified after Beccaluva
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56 1021 et al., (2009); other data from Baker et al. (1996), Pik et al. (1998), Kieffer et al. (2004),
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58 1022 Teklay et al. (2005) and Natali et al. (2011). LT = Low-Ti tholeiites, HT1 = High-Ti
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60 1023 tholeiites, HT2 = very High-Ti transitional basalts and picrites. Sampling locations are
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1024 indicated by circles (LT basalts), triangles (HT1 basalts), diamonds (HT2 basalts) and squares
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2 1025 (HT2 picrites). Miocene volcanoes overlapping CFBs are also reported. A paleogeographic
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4 1026 restoration at the Oligocene of the original Ethiopian-Yemeni plateau displaying a contiguous
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7 1027 zonal arrangement of LT, HT1 and HT2 lavas is reported in the inset.
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12 1029 **Figure 2** – Total Alkali-Silica (TAS) classification diagram for the Ethiopian-Yemeni CFBs.
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14 1030 The dashed line subdivides the alkaline and subalkaline series. LT = Low-Ti tholeiitic
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16 1031 basalts, HT1 = High-Ti tholeiitic basalts, HT2 = very High-Ti transitional basalts (diamonds)
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19 1032 and picrites (squares), MI = Melt Inclusions in olivine from HT2 picrite and basalt, Gl =
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21 1033 interstitial Glasses from picrite LAL73.
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26 1035 **Figure 3** – FeO_t, TiO₂, Nb and Ce vs MgO variation diagrams for Ethiopian-Yemeni CFB
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29 1036 (large symbols from Author data base; small symbols from Baker et al., 1996; Pik et al.,
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31 1037 1998; Teklay et al., 2005). Empirical boundaries between LT, HT1 and HT2 are drawn in
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33 1038 order to minimize the misclassified samples (generally less than 5%). Clear compositional
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35 1039 variations between magma types can be interpreted in terms of intergroup source enrichments
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38 1040 and fractionation trends. Symbols and other abbreviations as in Fig. 1 and 2.
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43 1042 **Figure 4** – Ch-Normalized REE patterns (a) and PM-Normalized incompatible elements
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45 1043 distribution (b) of HT2 lavas from Ethiopian-Yemeni CFB. The PM-Normalized distribution
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48 1044 of melt inclusions (MI) in olivine and interstitial glass (Gl) is reported in (c) and compared
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51 1045 with the compositional field of HT2 picrites (shaded area). Normalizing factors from Sun &
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53 1046 McDonough (1989). See text for further details.
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58 1048 **Figure 5** – Ch-Normalized REE patterns of clinopyroxenes from HT2 picrites (a) and basalts
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61 1049 (b) from the Ethiopian-Yemeni CFB. The same symbol is referred to different clinopyroxene
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1050 analyses from the same rock sample. Note that clinopyroxenes from the Yemeni HT2 basalt
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21051 (x), show the highest incompatible element patterns consistent with their comparatively more
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41052 alkaline character. Normalizing factors from Sun & McDonough (1989).
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101054 **Figure 6** – MgO vs FeO diagram for the Ethiopia-Yemeni CFB. HT2, HT1 and LT primary
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121055 magmas are modelled according to Herzerberg and Asimow (2015). Several liquid lines of
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141056 descent (Petrolog3, Danyushevsky and Plechov, 2011) satisfactorily match the observed
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171057 compositional variations; the HT2, HT1 and LT suites result from Ol+Cpx, Ol+Cpx+Pl and
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191058 Ol+Pl+Cpx fractionation trends, respectively (inset). For comparison, calculated primary
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221059 magmas and related compositional fields (shaded areas) of lavas from East Pacific Rise
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241060 (EPR), Iceland, Hawaii and Gorgona, together with equilibrium olivine arrays, are also
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271061 shown (modified after Herzberg et al., 2007 and references therein).
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321063 **Figure 7** – Fo_{Ol} vs $Mg\#_{bulk\ rock}$ for HT2 Ethiopian-Yemeni basalts and picrites. Tie lines join
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341064 multiple olivine composition recorded in the same rock sample. Note that most of the olivine
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361065 analyses in picrites show lower Fo contents than expected from equilibrium Kd_{Mg-Fe} (0.28-
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391066 0.32), as effect of olivine accumulation.
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441068 **Figure 8** – Temperature ($^{\circ}C$) vs MgO (wt%) petrogenetic grid for HT2 and LT suites based
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461069 on modelling described in Fig. 6. The hypothetical thermo-compositional evolution from
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491070 HT2 primary melts (dark grey field) to the most differentiated basalt is depicted by trend 1:
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511071 modelling matches the composition of the most forsteritic olivine observed in equilibrium
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541072 with the host picrite for pressure $< 1GPa$. In principle this imply that primary magmas (*ca.*
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561073 $1570^{\circ}C$) raised nearly-adiabatically up to the base of the crust (35-40 km), where the olivine
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581074 liquidus was reached at temperature of *ca.* $1510^{\circ}C$. Fractional crystallization process was
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611075 mainly controlled by olivine down to $1250^{\circ}C$ where cotectic high-Mg# clinopyroxene
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1076 formed, marking the transition from picritic to basaltic composition. For LT lavas (trend 2),
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2 1077 T_p is around 1430°C (light gray circle) and fractionation is controlled by olivine, followed by
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4 1078 plagioclase at ca. 1200°C. OIB and MORB data (open circles) are reported after Herzberg et
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7 1079 al. (2007) and references therein; T_p calculated for other CFB provinces (open squares) are
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9 1080 from this work.

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14 1082 **Figure 9** – PM-Normalized incompatible element distribution of primary magmas and
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16 evolved basalts for HT2 (a) and LT (b) suites. Fractional Crystallization (FC) modelling
17 1083 starting from HT2 and LT primary magmas satisfactorily matches the composition of the
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19 1084 observed differentiated basalt by removal of relative amounts of solid phases (Ol + Cpx for
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21 1085 HT2; Ol+ Pl + Cpx for LT) obtained by major element mass balance calculation.

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29 1088 **Figure 10** – (a) PM-Normalized trace element distribution for HT2 picrites from the
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31 1089 Ethiopian-Yemeni CFB (shaded field), compared with high-MgO lavas from other CFB and
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33 OIB provinces. Data after Gaffney (2002); Gibson et al. (2000 and references therein); Kerr
34 1090 et al. (1996); Kamenetsky et al. (2012); Révillon et al. (2000); Thompson et al. (2001); Arndt
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36 1091 et al. (1995, 1998); Carlson et al. (2006); Heinonen and Luttinen (2008 and references
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38 1092 therein); Heinonen et al. (2010); Ellam and Cox, (1989, 1991); Melluso et al. (1995). (b)
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41 1093 Incompatible element distribution of calculated primary HT2 picrites (shaded field) and of
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43 1094 theoretical melts obtained by batch ($F = 0.10$) and fractional ($F = 0.09$) melting of a mantle
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45 1095 source composed by 85% of lherzolite (xenolith GOJ40A; Beccaluva et al., 2009) and 15% of
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47 1096 eclogite (included in Hoggar Neoproterozoic terranes; Berger et al., 2014). Source mode and
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49 1097 melting proportions conform to experimental data by Walter (1998); partition coefficients
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51 1098 (K_d) from the GERM database. Normalizing factors from Sun & McDonough (1989).

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1101 **Figure 11** – Sr-Nd-Pb isotopic composition of Ethiopian-Yemeni lavas including data from
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21102 Pik et al. (1999) and Teklay et al. (2005). (a) Sr-Nd isotopic composition (expressed as $\epsilon(t)$)
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41103 of the Ethiopian-Yemeni Oligocene lavas in comparison with those recorded in other
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71104 Phanerozoic CFB provinces worldwide (data from Gibson et al. 2000; Iacumin et al., 2003;
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91105 Heinonen et al., 2014 and references therein; Sheth et al. 2014). (b) $^{207}\text{Pb}/^{204}\text{Pb}$ vs $^{206}\text{Pb}/^{204}\text{Pb}$
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121106 of the Ethiopian-Yemeni lavas compared with other Phanerozoic plateau basalts, MORB,
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141107 OIB, geochrons and Northern Hemisphere Reference lines reported by Jackson and Carlson
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171108 (2011, and references therein).

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221110 **Figure 12** – $^{87}\text{Sr}/^{86}\text{Sr}$ - $^{143}\text{Nd}/^{144}\text{Nd}$ (a), $^{208}\text{Pb}/^{204}\text{Pb}$ - $^{206}\text{Pb}/^{204}\text{Pb}$ (b) and $^{206}\text{Pb}/^{204}\text{Pb}$ - $^{87}\text{Sr}/^{86}\text{Sr}$ (c)
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241111 diagrams for the Ethiopian-Yemeni CFB. Large symbols refer to data from this work,
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271112 whereas small symbols to data from Pik et al. (1999) and Teklay et al. (2005). Notional
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291113 mantle end-members (DM, HIMU, EMI, EMII) are after Zindler and Hart (1986) and Stracke
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311114 et al. (2005). The $^{206}\text{Pb}/^{204}\text{Pb}$ - $^{87}\text{Sr}/^{86}\text{Sr}$ isotopic variation observed in cpx-bulk rock pairs is
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341115 shown in (d), where isotopic ratios have been recalculated for initial values at 30 Ma.

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391117 **Figure 13** – TiO_2 (wt%) vs $^{87}\text{Sr}/^{86}\text{Sr}$ (a) and TiO_2 (wt%) vs $^{206}\text{Pb}/^{204}\text{Pb}$ (b) diagrams for the
40
411118 Ethiopian-Yemeni CFB. Large symbols refer to data from this work, whereas small symbols
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441119 to data from Pik et al. (1999) and Teklay et al. (2005). Note that isotopic composition of
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461120 Ethiopian CFB show a good correlation with TiO_2 content, whereas the Yemeni samples are
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48
491121 displaced toward comparatively lower $^{87}\text{Sr}/^{86}\text{Sr}$ and higher $^{206}\text{Pb}/^{204}\text{Pb}$ isotopic values (see
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511122 text for explanation).

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531123
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561124 **Figure 14** – Schematic model for the generation of Oligocene Ethiopian-Yemeni CFB by the
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581125 Afar plume (modified after Beccaluva et al., 2009 and updated with data from this work). The
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60
611126 impingement of the Afar plume on the Afro-Arabian lithosphere caused regional doming and

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1127 a dramatic thermal anomaly (excess temperature T_{ex} , 300-350°C) that triggered extensive
1
21128 melting and eruption of CFB from the plume head. Thermo-chemical effects in the plume
3
41129 head progressively decreased outward as reflected by the CFB zonal arrangement with the
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6
71130 hottest HT2 picrites (T_p 1570°C) at the core and LT basalts (T_p 1430°C) at the periphery. T_{ex}
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9
101131 is the maximum temperature difference with respect to the ambient mantle whose
11
121132 temperatures are based on thermobarometric estimates of mantle xenoliths from the East-
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141133 African-Arabian domain (Conticelli et al., 1998; Beccaluva et al., 2011b, Sgualdo et al.,
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16
171134 2015). Variable amounts of eclogitic components entrained by the plume are required in the
18
191135 HT magma sources (up to 15% for HT2).
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21
221136 Seismic tomography data record a low-velocity anomaly in the upper mantle - extending for
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241137 at least 1000 km from northern Ethiopia to Yemen up to *ca.* 150 km depth – that is related to
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26
271138 the persisting effects of the Afar plume (Davaille et al., 2005; Chang and Van der Lee, 2011).
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291139 See text for further explanation.
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Figure 1

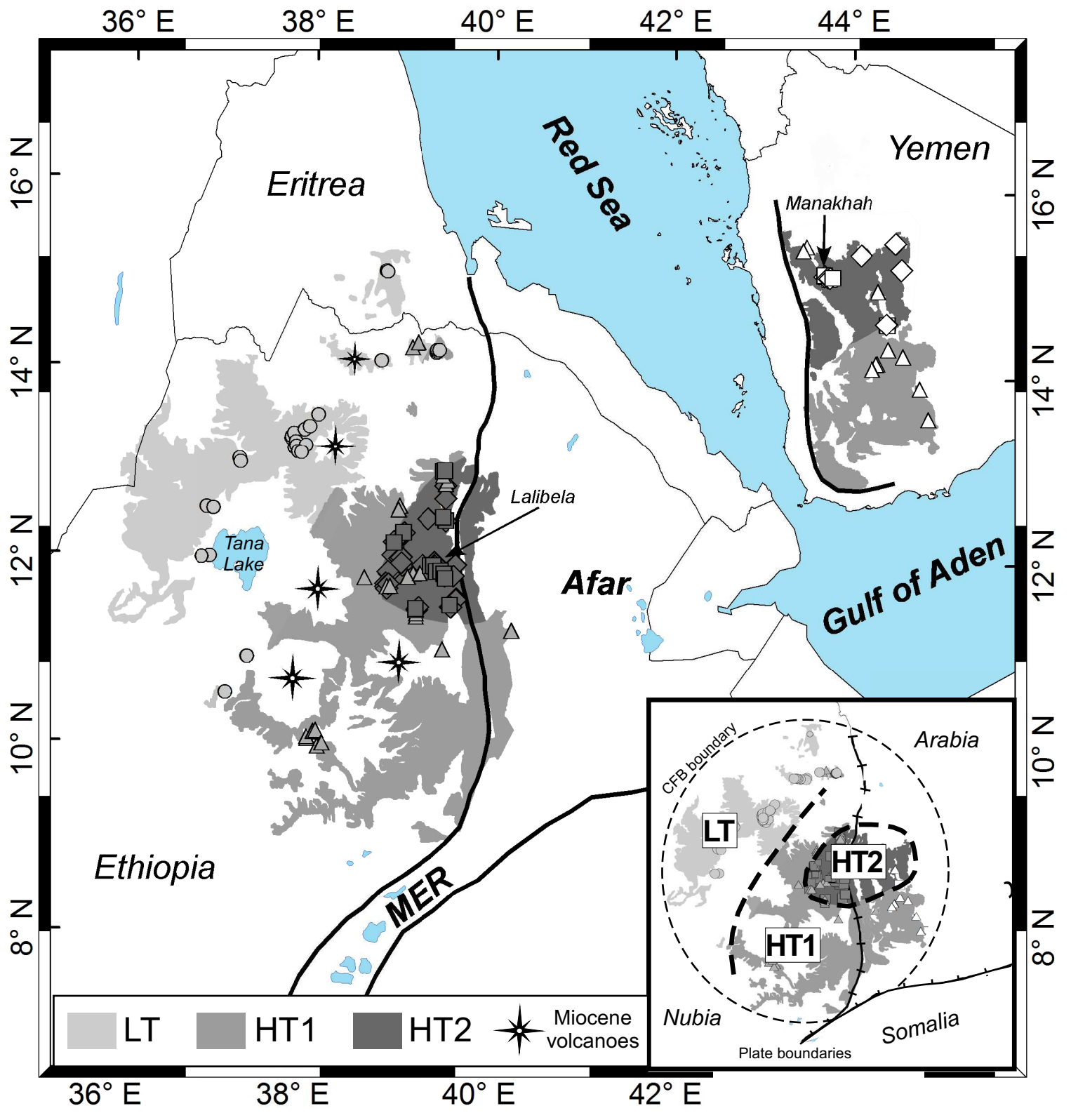
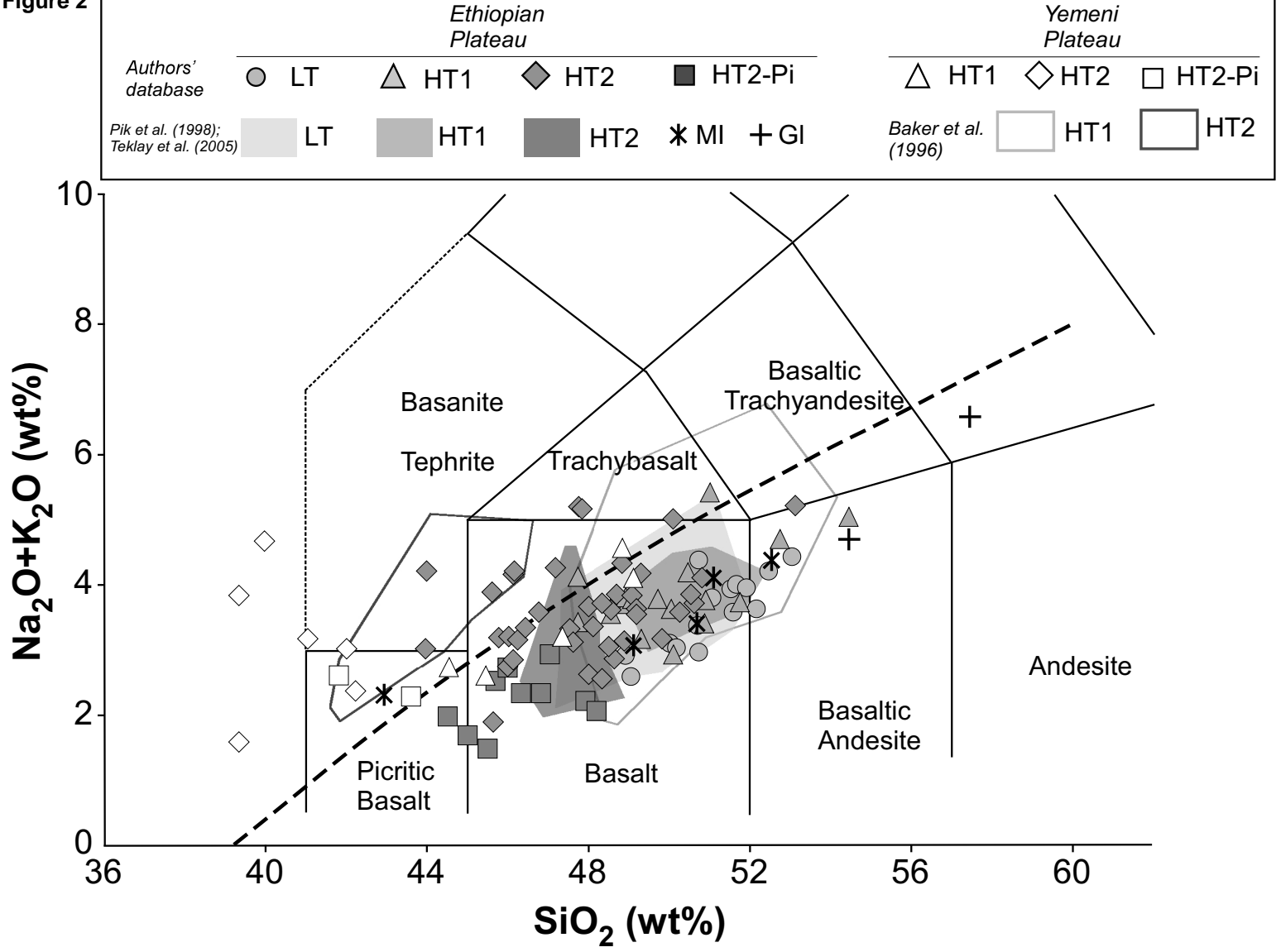


Figure 2



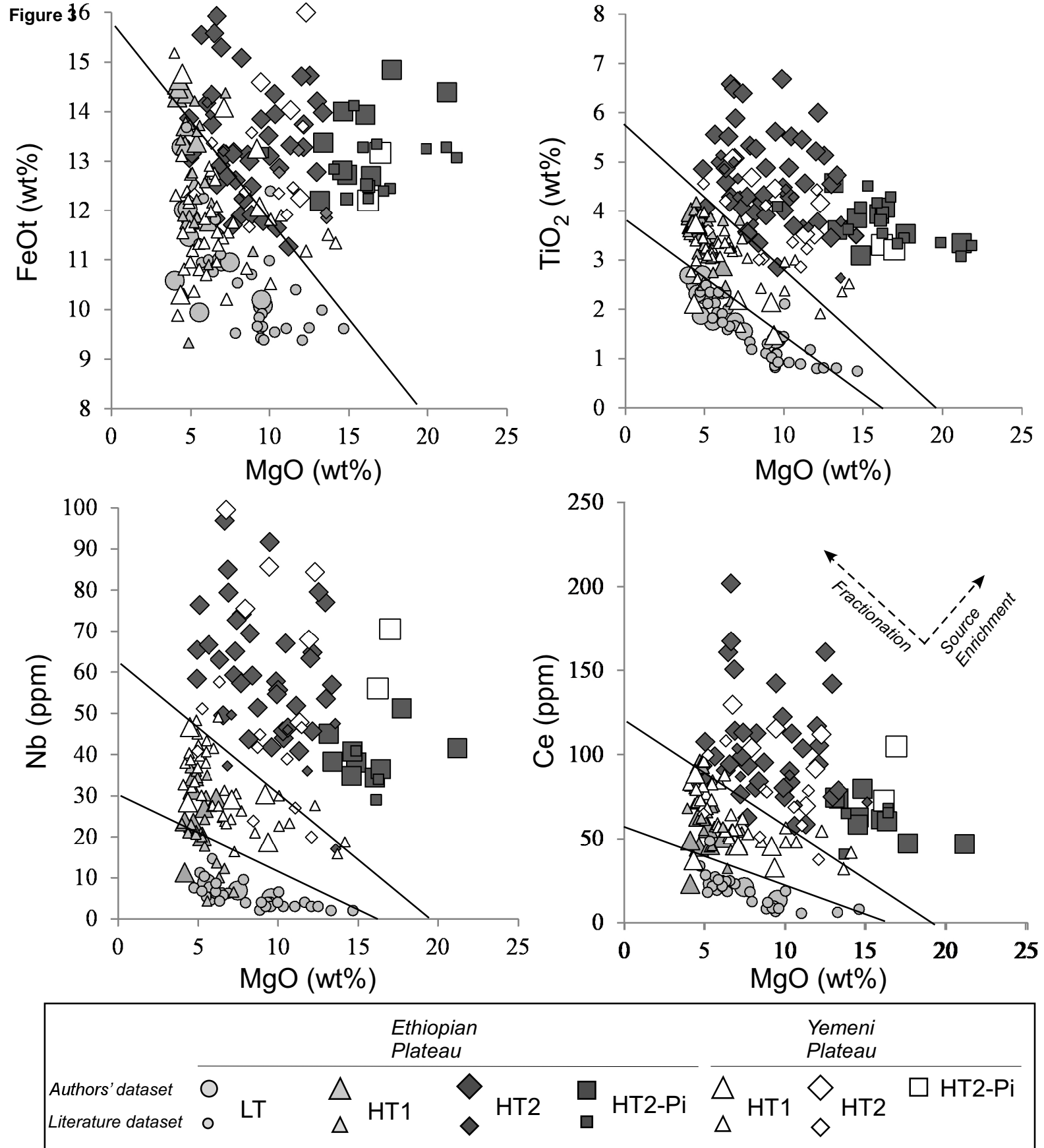


Figure 4

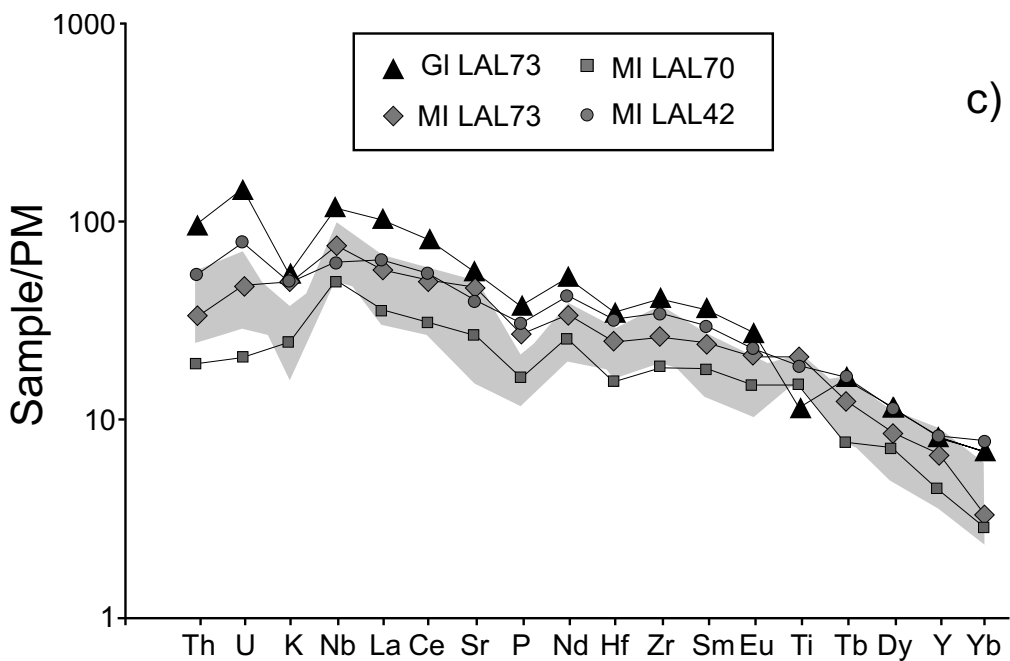
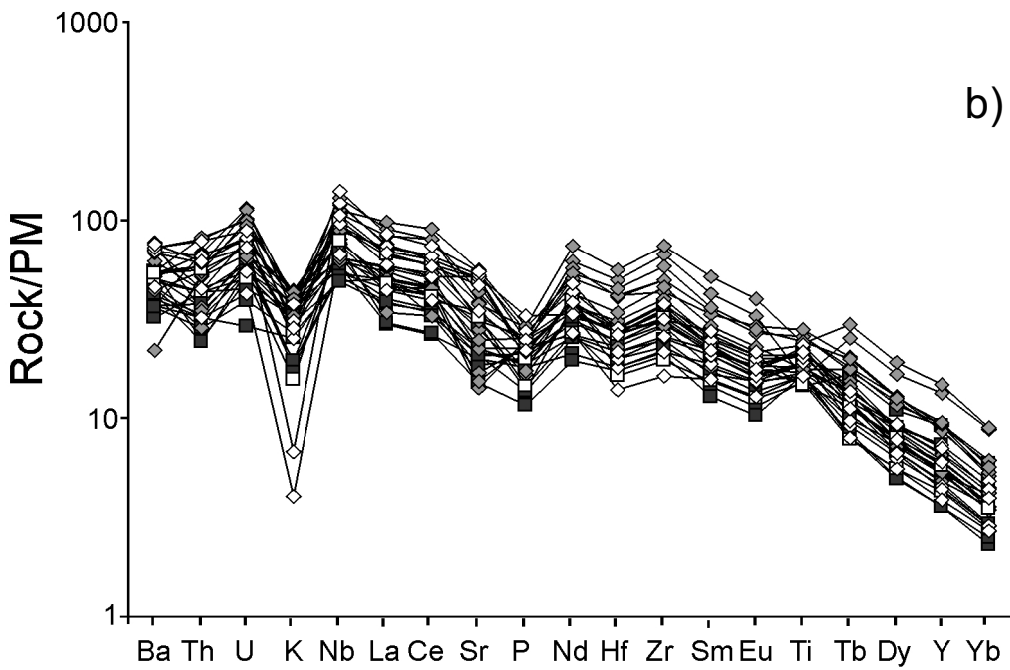
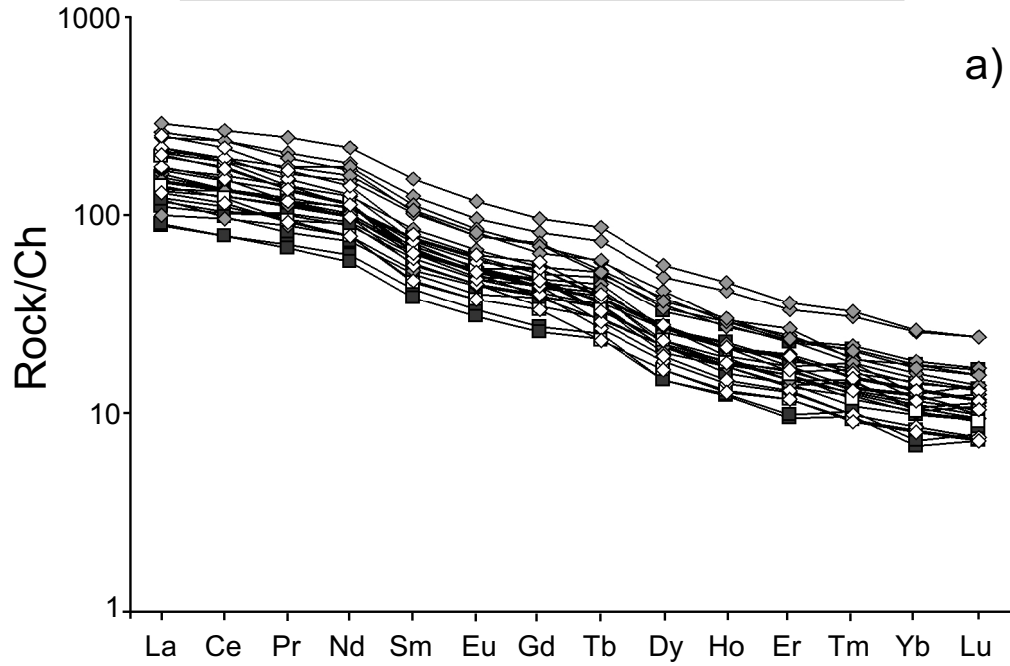
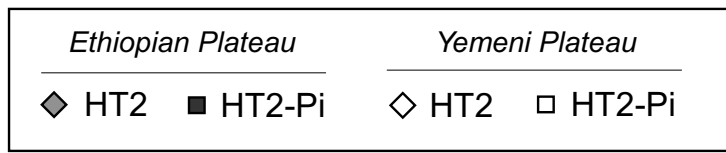


Figure 5

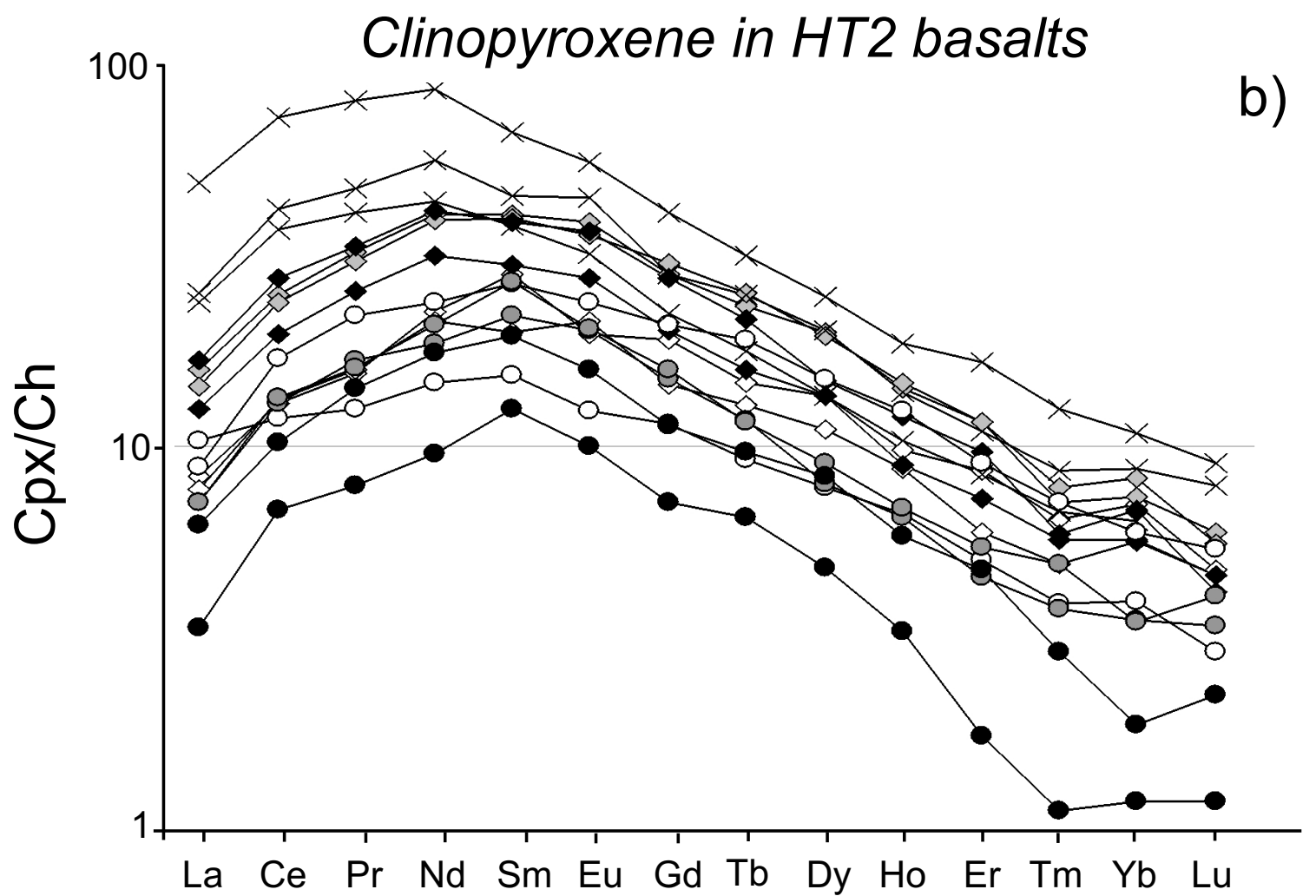
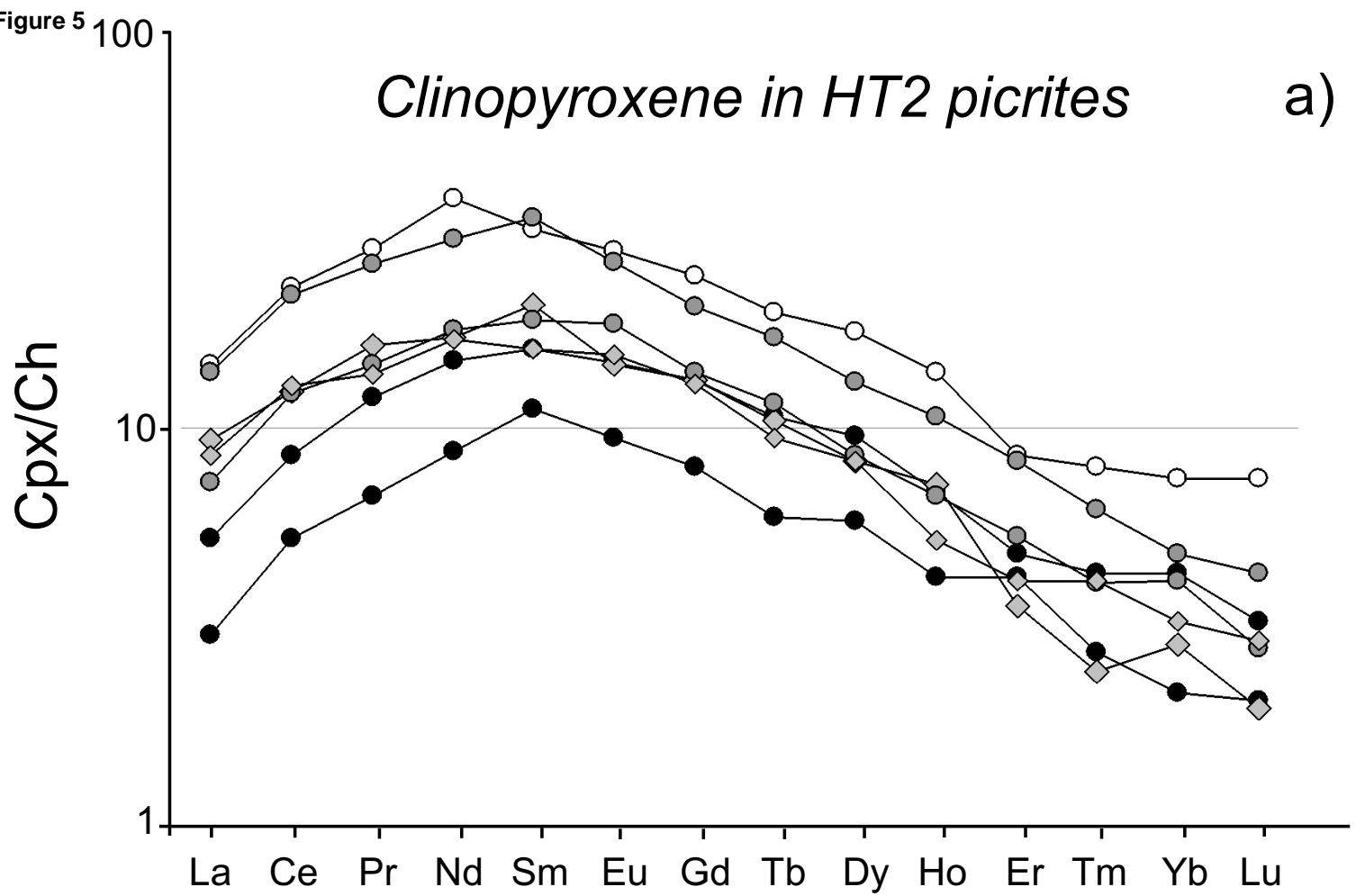


Figure 6

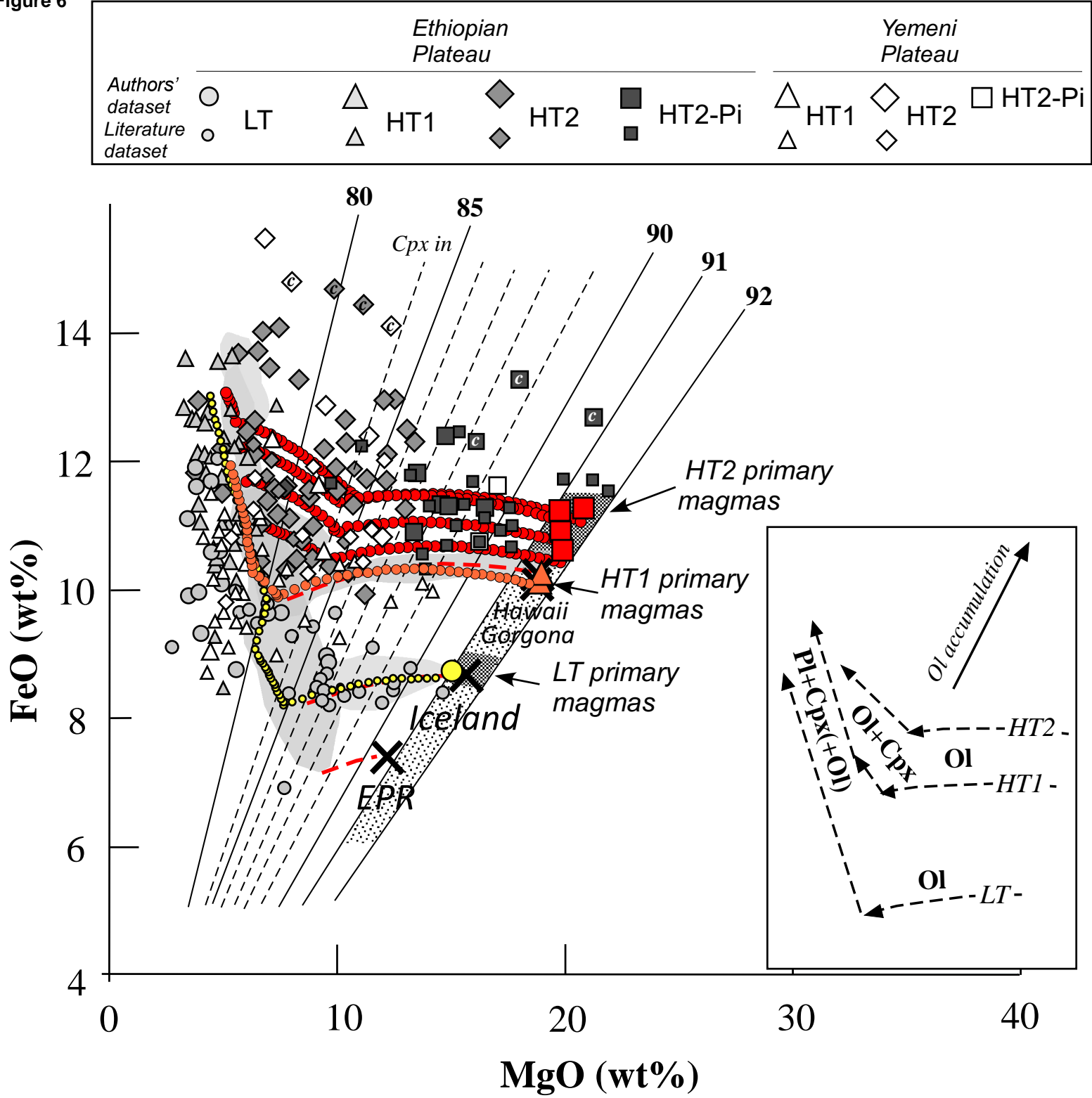


Figure 7

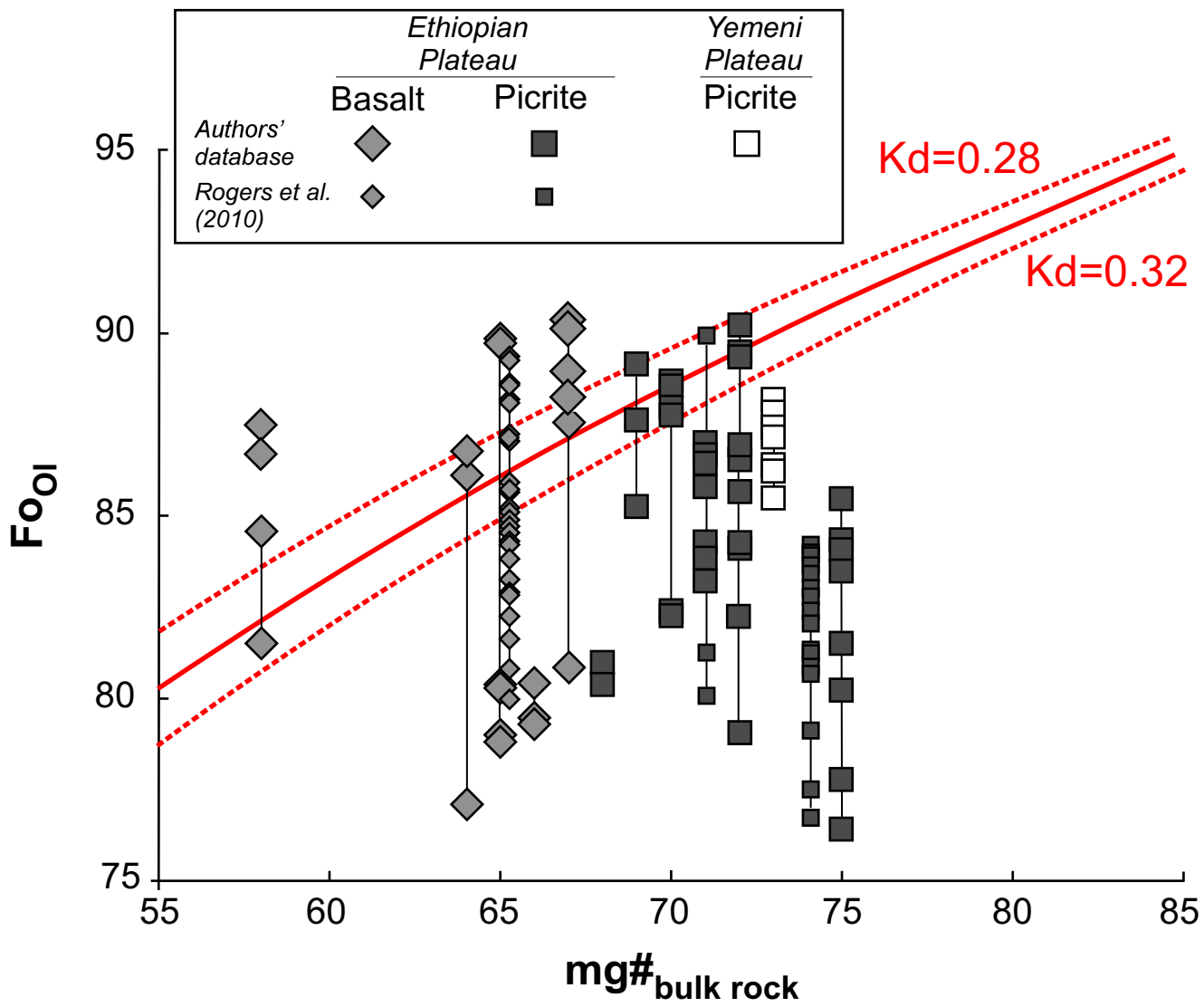


Figure 8

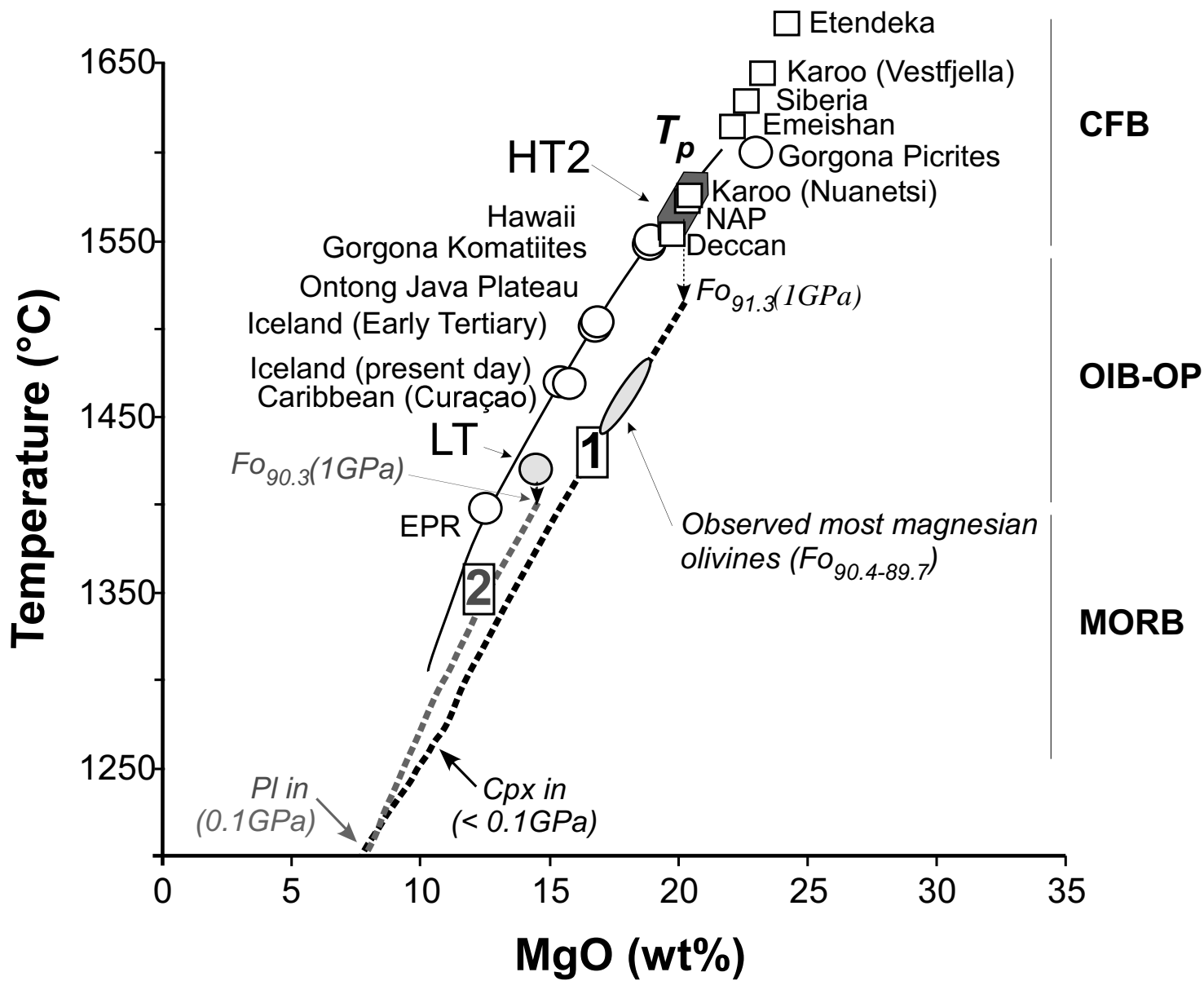


Figure 9

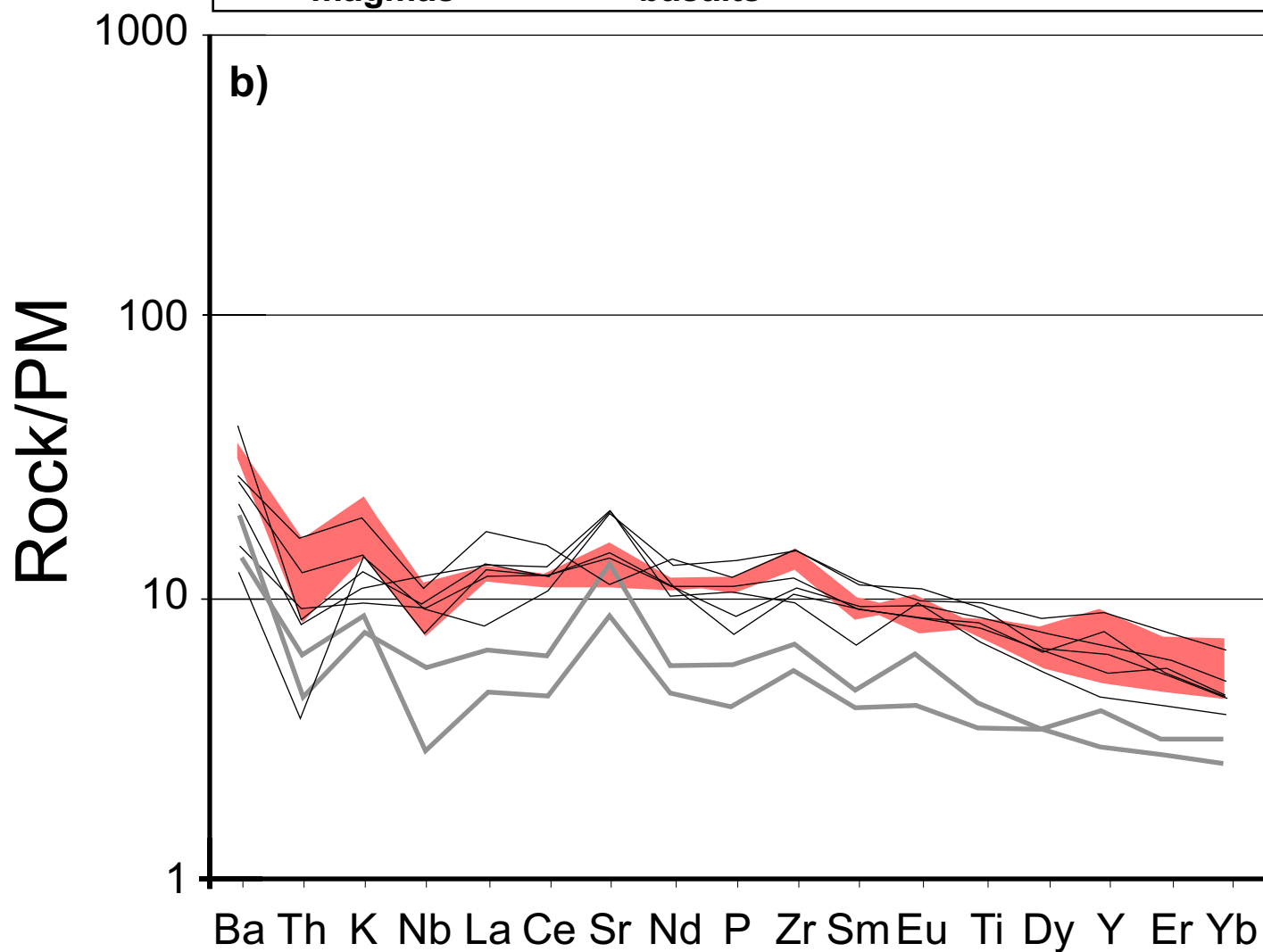
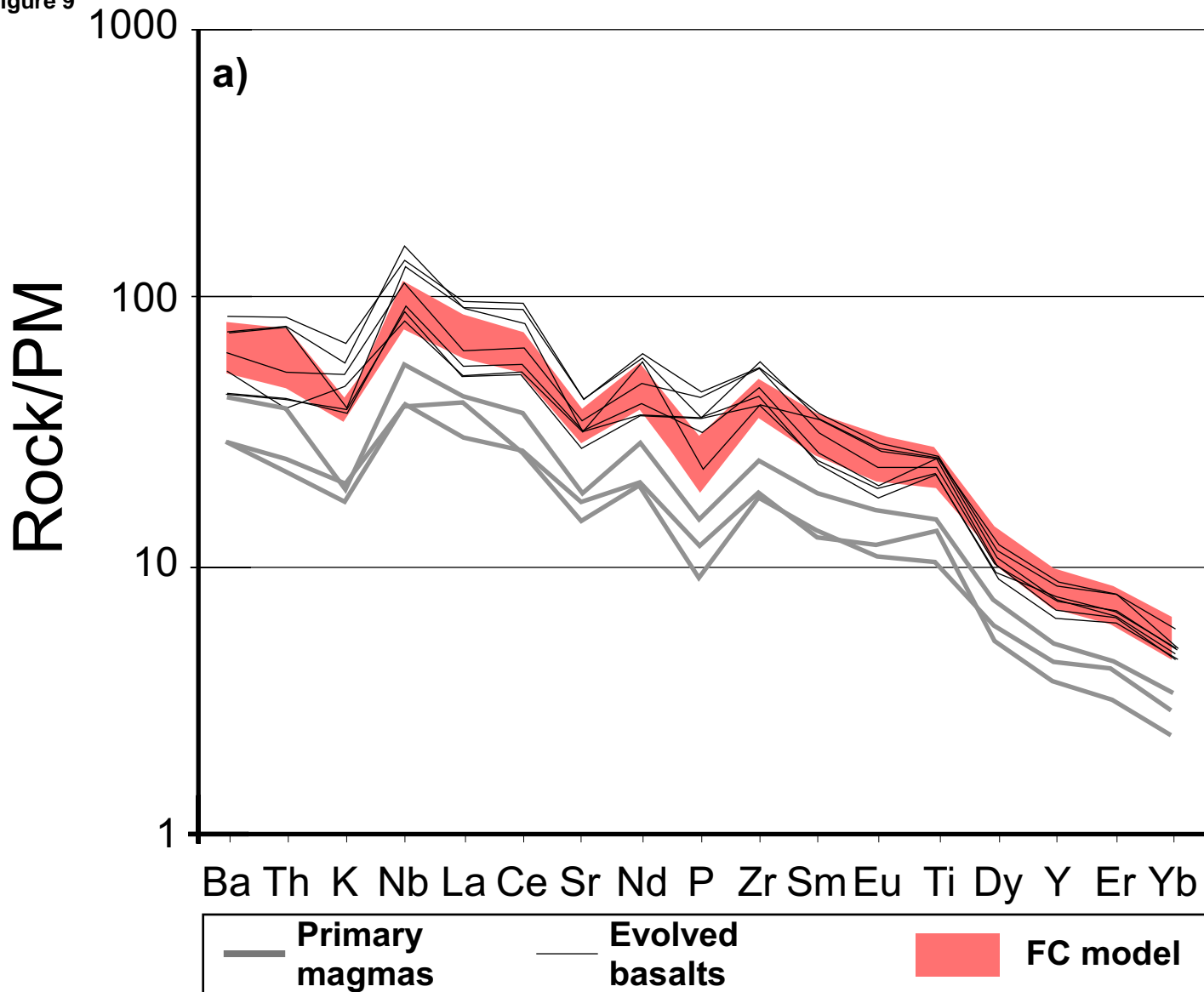
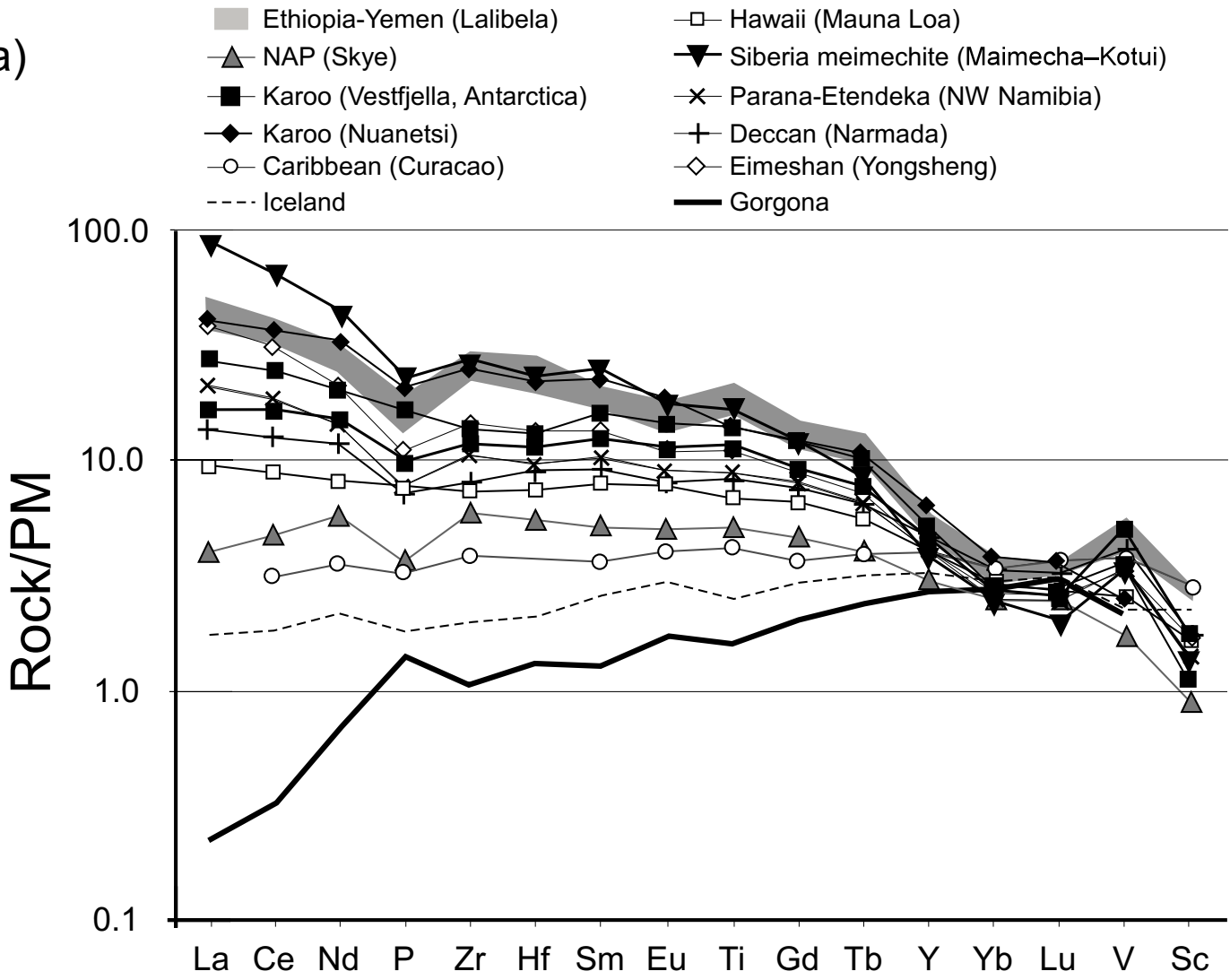


Figure 10

High MgO-lavas (13-22% MgO)

a)



b)

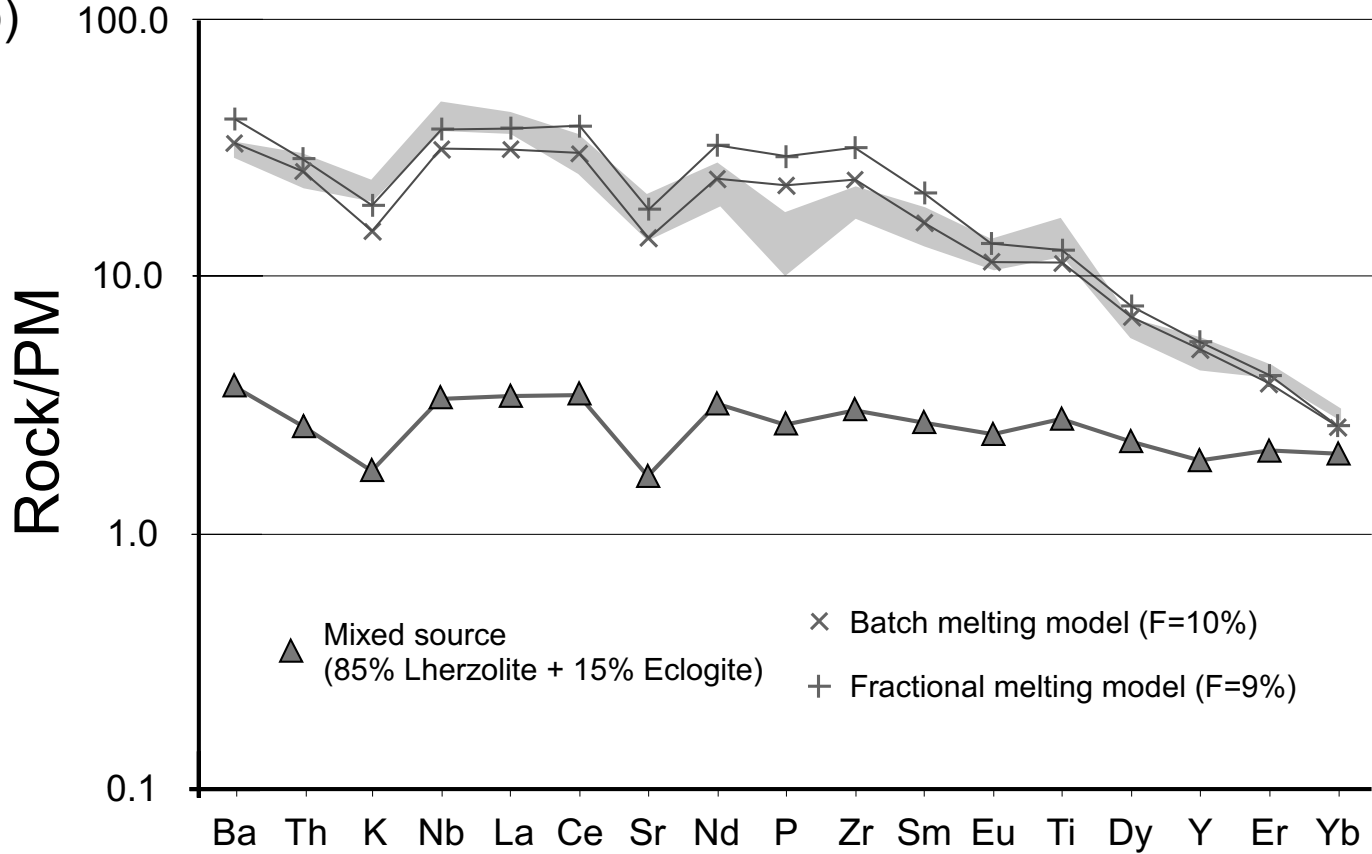
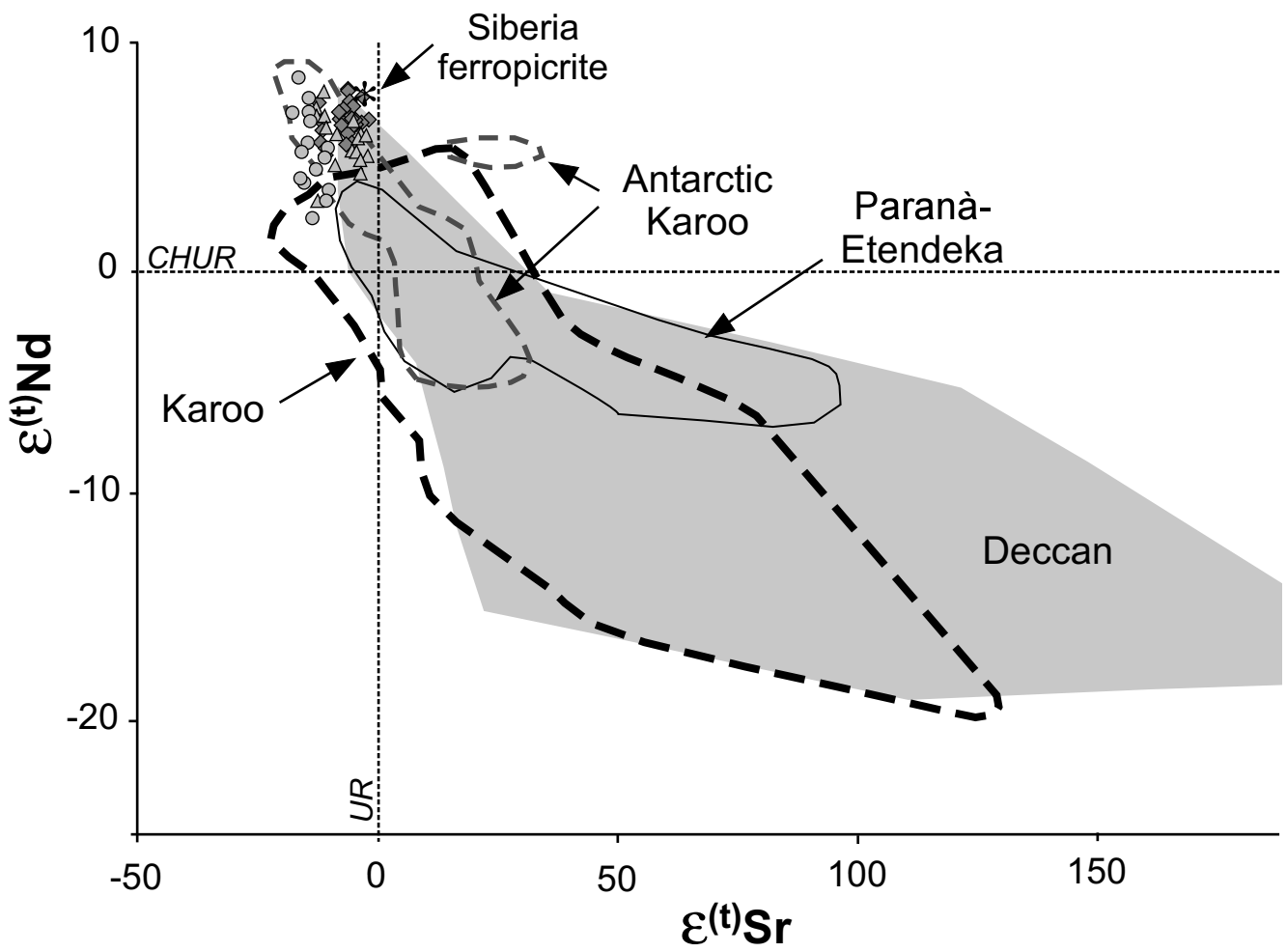


Figure 11

a)



b)

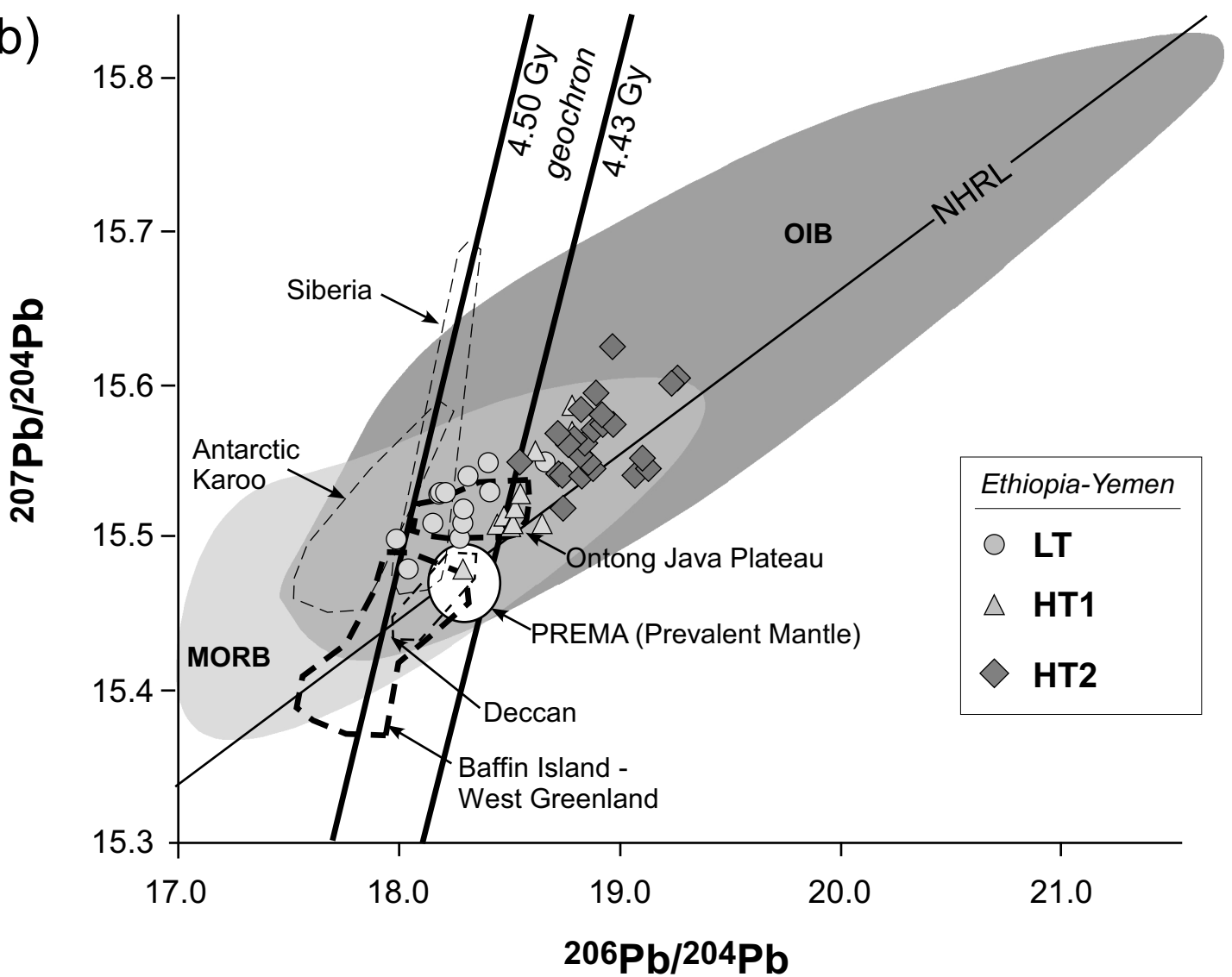


Figure 12

Ethiopian Plateau				Yemeni Plateau			
○	△	◆	■	◇	□		
○	△	◆	■	◇	□		
LT	HT1	HT2	HT2-Pi	HT2	HT2-Pi		

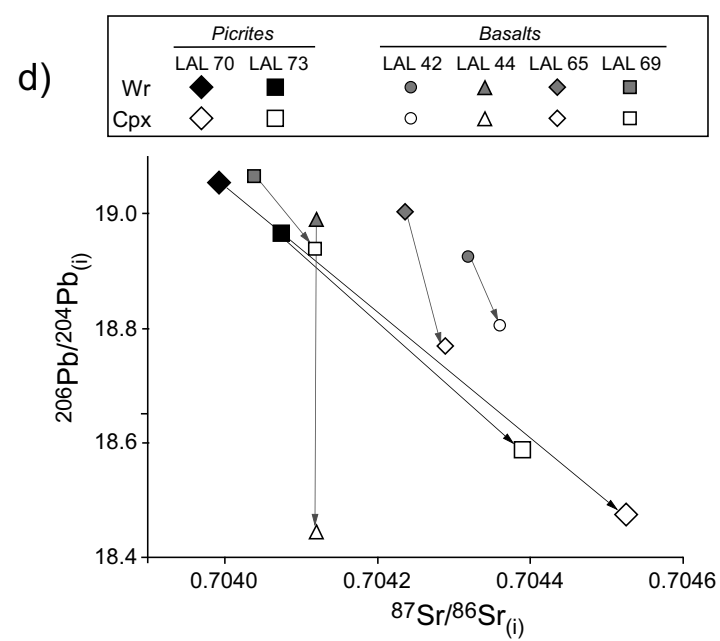
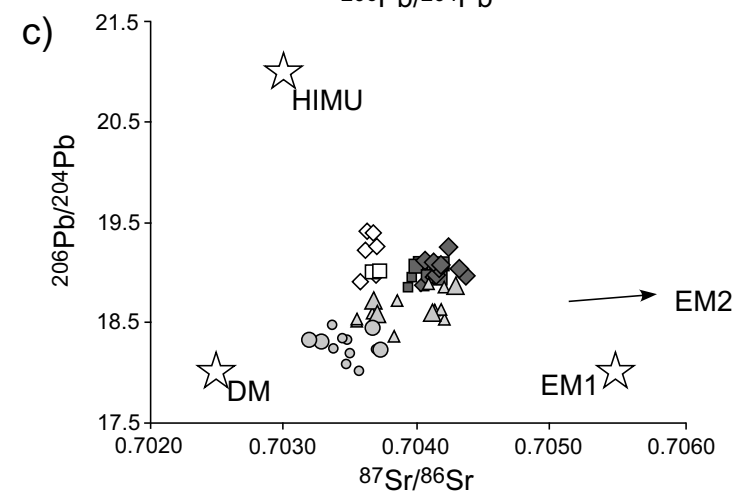
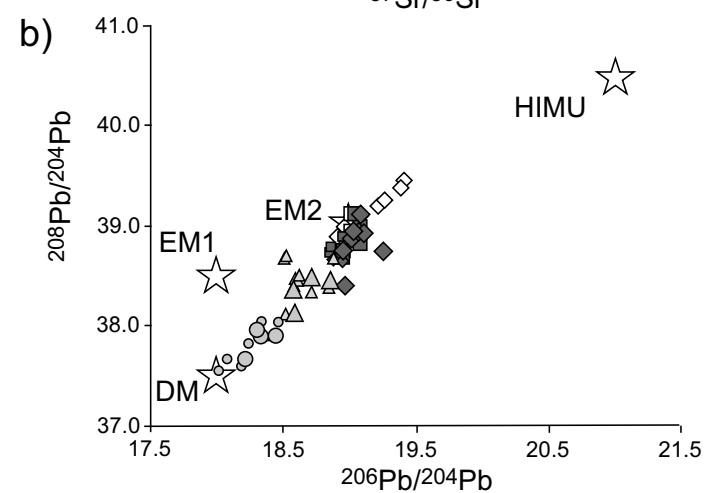
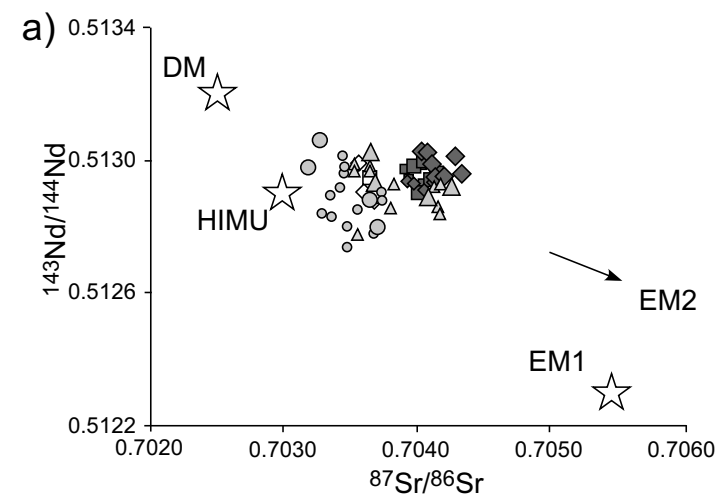


Figure 13

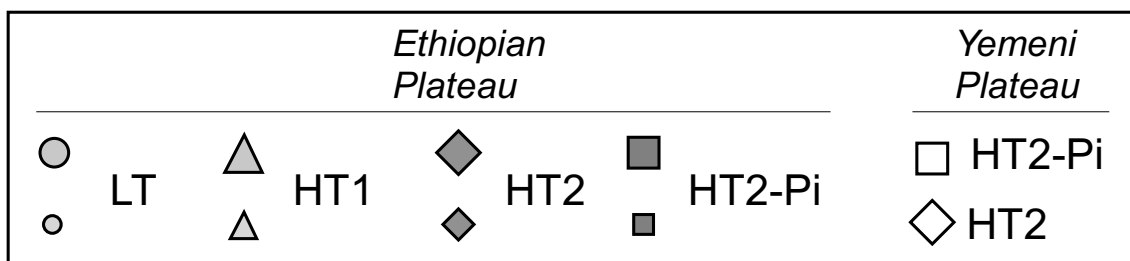
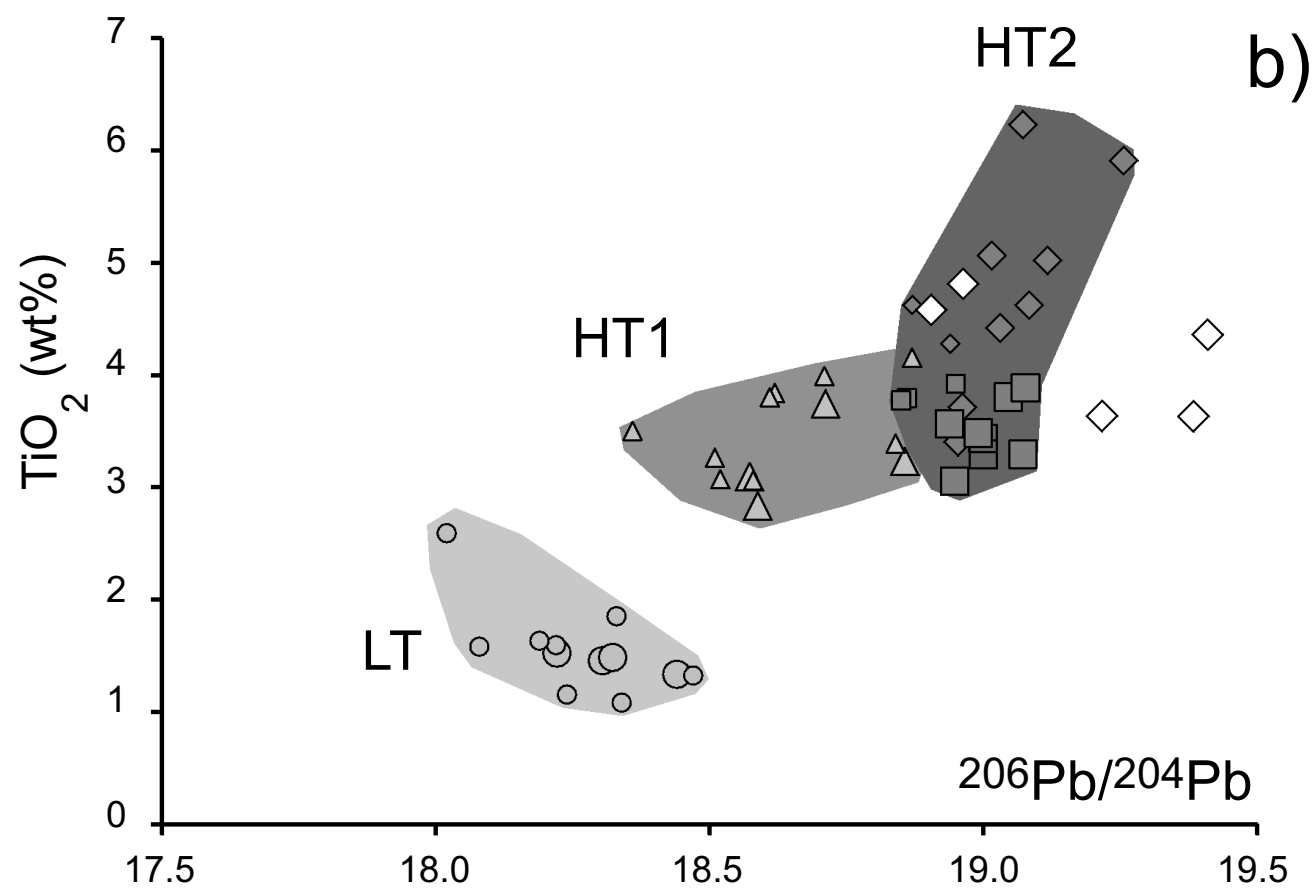
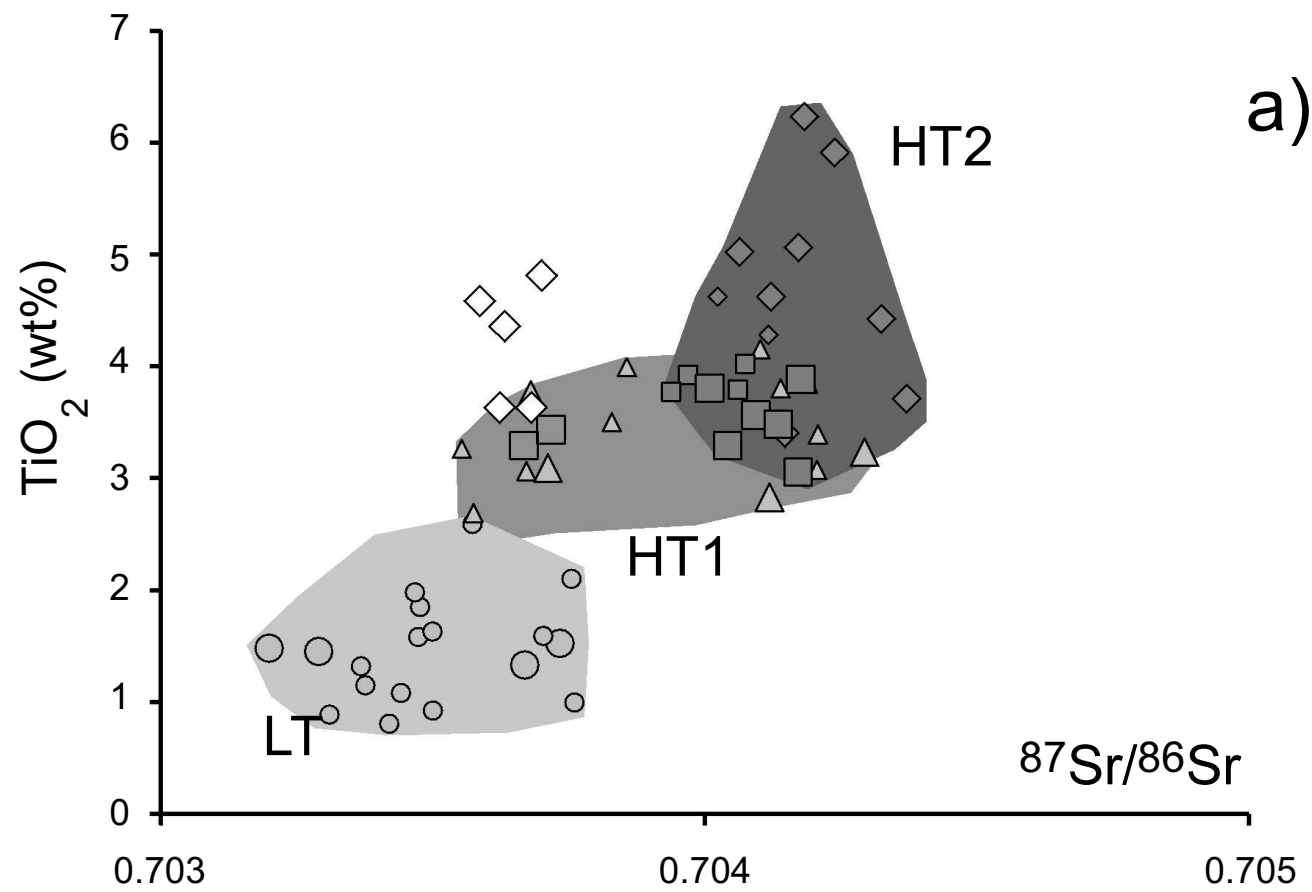


Figure 14

NW

SE

Ethiopia ← Future Afar Margin → Yemen

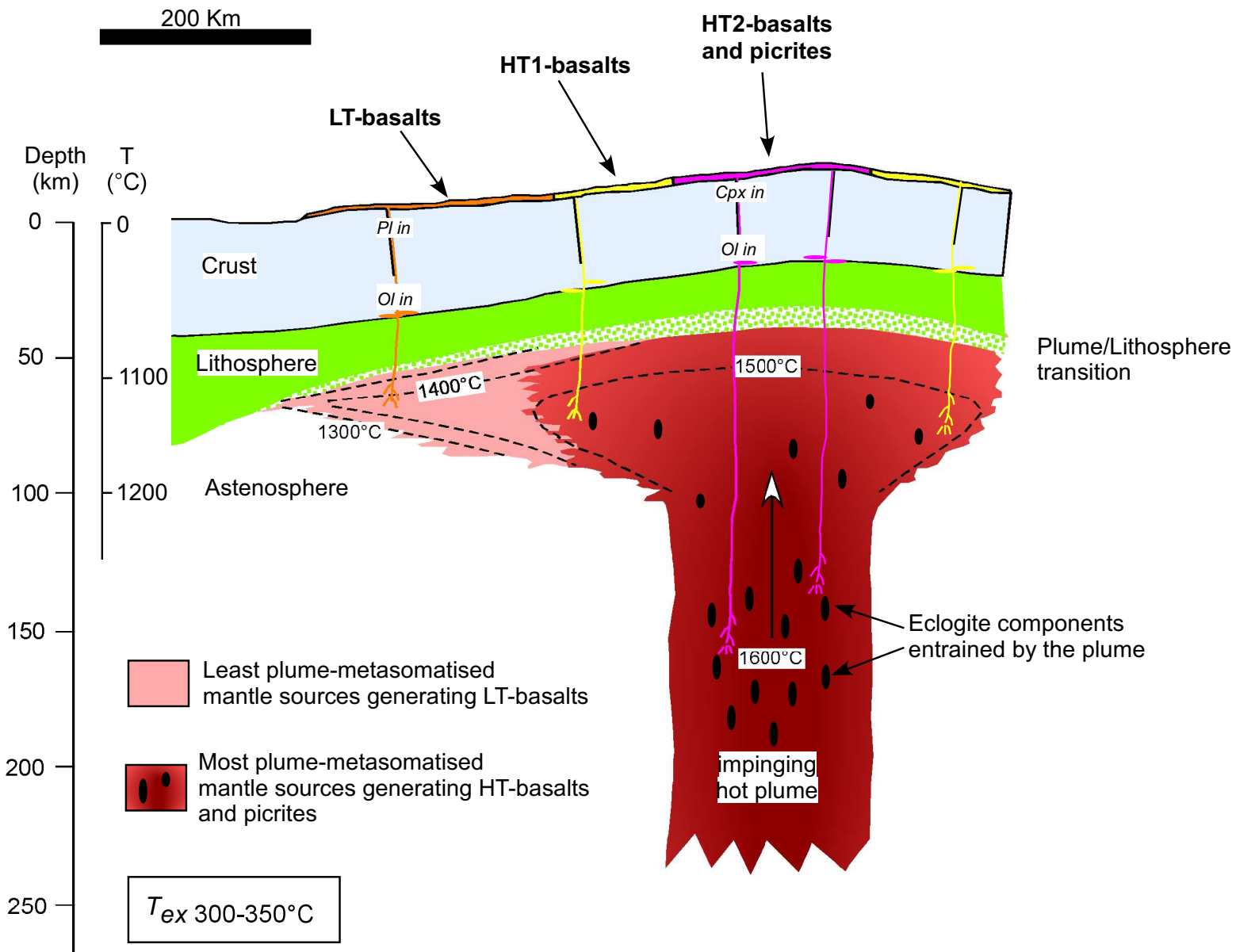


Table 1

Table 1

Ethiopian plateau (Lalibela suite)

Rock type	Picrite										Basalt											
	39.4200	39.4056	39.3950	39.2403	39.5544	39.4617	39.4308	39.5433	39.5217	39.5156	39.4275	39.0039	38.9861	39.2375	39.5514	39.5411	39.5003	39.0100	39.5478	39.5431	39.5297	39.5469
Lon. (dd)	39.4200	39.4056	39.3950	39.2403	39.5544	39.4617	39.4308	39.5433	39.5217	39.5156	39.4275	39.0039	38.9861	39.2375	39.5514	39.5411	39.5003	39.0100	39.5478	39.5431	39.5297	39.5469
Lat.	11.9664	11.9536	11.9442	11.4878	11.8242	11.8956	11.9464	12.4431	12.4742	11.8869	11.9742	11.9450	11.9336	11.4781	11.8458	11.8706	11.8892	12.2181	12.4350	12.4428	12.4561	12.8128
Sample	LAL5*	LAL6*	LAL9*	LAL33	LAL41	LAL45	LAL47*	LAL70	LAL73*	LAL1*	LAL3*	LAL14*	LAL15*	LAL31	LAL42	LAL43	LAL44	LAL60*	LAL65	LAL69	LAL72*	LAL76*
SiO ₂ (wt%)	46.57	45.66	46.83	47.79	48.15	45.77	46.26	44.52	44.36	48.06	50.60	46.26	45.84	48.55	48.14	45.64	45.41	47.21	45.80	45.75	48.12	48.92
TiO ₂	4.57	4.07	3.49	3.61	3.09	3.86	3.93	3.34	3.53	5.50	3.97	5.99	5.57	3.44	3.76	5.09	5.13	4.85	4.48	4.68	4.44	3.69
Al ₂ O ₃	8.69	8.40	8.87	8.99	8.30	8.62	7.79	6.02	6.87	8.14	11.12	7.88	9.34	9.22	9.57	7.95	8.35	8.51	8.10	7.48	8.87	9.99
FeO	10.75	11.19	11.29	11.78	11.21	12.34	12.28	12.68	13.07	10.23	10.47	12.09	12.10	11.22	11.63	12.87	12.76	12.55	12.31	12.19	11.79	11.32
Fe ₂ O ₃	1.61	1.68	1.69	1.77	1.68	1.85	1.84	1.90	1.96	1.53	1.57	1.81	1.81	1.68	1.74	1.93	1.91	1.88	1.85	1.83	1.77	1.70
MnO	0.20	0.18	0.18	0.18	0.18	0.18	0.17	0.18	0.18	0.17	0.17	0.19	0.19	0.18	0.17	0.17	0.16	0.17	0.16	0.17	0.17	0.17
MgO	13.16	16.40	14.59	13.39	14.89	14.64	16.04	21.21	17.70	10.43	8.17	12.13	9.40	12.92	12.06	12.44	11.83	10.24	12.78	13.22	9.84	10.61
CaO	10.15	9.37	10.41	9.58	9.97	9.15	8.74	7.79	8.89	11.81	9.85	9.96	10.46	9.20	9.57	9.94	9.35	10.19	9.26	10.48	10.62	9.97
Na ₂ O	1.92	1.75	1.55	1.65	1.53	1.63	1.44	0.98	1.10	2.02	2.43	1.98	2.95	1.84	1.69	1.53	1.88	2.15	2.11	1.87	1.84	2.06
K ₂ O	0.97	0.75	0.80	0.57	0.54	1.10	0.89	0.98	0.57	1.31	1.28	1.18	1.19	1.01	0.85	1.19	1.27	1.16	1.17	0.96	1.20	1.09
P ₂ O ₅	0.41	0.40	0.29	0.40	0.30	0.46	0.44	0.39	0.25	0.61	0.36	0.53	0.50	0.53	0.37	0.63	0.55	0.48	0.52	0.50	0.55	0.49
LOI	1.01	0.16	0.00	0.30	0.16	0.42	0.17	0.00	1.51	0.16	0.00	0.02	0.65	0.20	0.45	0.63	1.40	0.62	1.46	0.85	0.80	0.00
mg#	0.69	0.72	0.70	0.67	0.70	0.68	0.70	0.75	0.71	0.65	0.58	0.64	0.58	0.67	0.65	0.63	0.62	0.59	0.65	0.66	0.60	0.63
Ni (ppm)	646	1097	697	509	569	565	724	984	836	237	233	611	214	481	439	318	385	168	434	411	212	327
Co	67	71	71	76	80	77	92	106	92	45	48	59	56	73	69	76	74	64	71	73	66	65
Cr	808	1551	1203	899	1348	995	1100	1820	1240	595	494	814	478	859	817	625	583	433	976	1022	528	723
V	449	371	343	360	345	345	343	305	356	475	400	482	513	343	353	386	399	442	367	390	383	361
Sc	44	43	43	40	41	37	38	36	33	53	41	50	43	40	38	51	43	46	38	45	48	38
Rb	15	5	7	15	28	26	19	22	24	15	20	19	20	23	21	23	26	28	28	31	20	23
Sr	457	470	392	435	396	460	442	555	454	596	516	684	673	407	496	678	608	672	389	457	570	471
Ba	336	259	253	258	225	288	276	383	292	377	426	369	507	258	305	385	365	428	359	535	381	310
Zr	309	265	249	315	240	342	356	288	282	436	320	495	442	307	329	499	510	478	428	434	406	316
Hf	8.72	6.68	6.15	8.26	8.63	7.40	7.54	5.53	5.30	12.8	8.17	12.9	12.5	15.3	10.5	17.3	13.9	8.83	8.31	8.23	9.45	7.06
Nb	45.0	36.5	34.8	38.2	38.1	40.8	34.4	41.5	51.2	67.0	43.7	64.8	91.7	77.0	45.6	79.5	63.4	45.6	53.5	56.9	54.6	46.5
Ta	2.66	2.10	2.11	1.38	1.39	1.43	1.26	2.41	2.93	4.03	2.68	4.39	7.10	2.61	1.65	2.79	2.28	1.66	3.10	3.34	3.27	2.72
Th	3.68	2.72	2.41	3.17	3.67	2.80	2.72	2.08	2.05	4.37	4.48	5.71	6.73	5.78	4.61	6.95	4.93	3.74	3.06	2.92	4.92	2.43
U	0.93	0.60	0.81	1.30	1.28	1.25	1.13	1.03	1.14	1.39	1.59	1.64	2.11	2.42	2.08	2.10	1.97	1.68	1.38	1.18	2.35	1.32
Y	26.4	21.4	25.1	33.2	41.3	23.6	24.7	16.1	16.2	41.8	30.0	39.2	39.7	60.1	42.0	67.3	42.8	52.6	28.0	24.4	50.1	41.6
La	33.2	26.1	35.0**	30.8	34.0	27.3	25.7	20.4	20.8	47.5	37.7	45.2	61.2	57.5	40.0	67.4	49.2	39.0	30.7	31.2	34.5	23.3
Ce	74.5	60.3	58.3	74.0	79.6	62.7	61.5	46.7	47.1	113	80.5	105	142	142	94.4	161	117	90.2	75.0	78.9	80.4	57.7
Pr	11.0	8.8	7.57	10.3	11.3	9.5	9.2	6.6	6.3	15.16	10.74	16.2	18.2	19.2	13.2	22.9	16.4	12.3	10.5	10.8	11.4	8.20
Nd	43.5	35.0	33.9	45.0	49.7	42.0	40.9	28.7	26.4	68.9	45.1	68.0	77.6	83.5	57.2	101	72.7	51.4	46.2	46.7	50.1	35.7
Sm	9.30	7.40	7.63	10.0	11.6	9.5	9.1	6.3	5.7	15.3	9.64	14.7	15.5	18.6	12.9	23.0	16.1	10.3	10.5	10.1	11.2	7.96
Eu	3.05	2.55	2.27	2.98	3.48	3.05	2.69	1.90	1.74	4.45	2.79	4.83	4.50	5.47	3.81	6.64	4.66	2.89	3.13	2.96	3.44	2.49
Gd	8.73	6.93	7.65	8.94	10.3	10.8	8.15	5.52	5.18	14.6	9.52	13.8	14.5	16.4	11.1	19.3	13.1	7.91	9.27	8.72	9.61	7.86
Tb	1.40	1.13	1.07	1.49	1.78	1.44	1.29	0.91	0.87	1.95	1.31	2.18	1.85	2.72	1.91	3.20	2.16	1.26	1.56	1.43	1.64	1.24
Dy	6.23	5.00	5.47	6.63	8.11	5.57	5.37	3.68	3.61	9.46	6.62	9.43	8.87	12.1	8.51	13.9	9.23	5.30	6.34	5.74	6.64	5.01
Ho	1.11	0.90	0.97	1.25	1.54	1.01	1.02	0.68	0.69	1.59	1.17	1.69	1.50	2.29	1.60	2.54	1.67	0.95	1.19	1.06	1.24	0.96
Er	2.40	1.94	2.51	2.97	3.68	2.61	2.39	1.54	1.59	4.01	3.06	3.68	3.79	5.44	3.80	5.87	3.85	2.21	2.64	2.36	2.77	2.25
Tm	0.38	0.30	0.31	0.42	0.52	0.33	0.32	0.24	0.25	0.47	0.38	0.52	0.44	0.77	0.54	0.81	0.52	0.31	0.42	0.36	0.45	0.36
Yb	1.84	1.46	1.79	2.36	2.93	2.03	1.83	1.14	1.22	2.64	2.19	2.63	2.50	4.30	3.01	4.39	2.81	1.68	1.97	1.71	2.09	1.77
Lu	0.26	0.25	0.24	0.33	0.41	0.26	0.24	0.18	0.19	0.34	0.29	0.42	0.33	0.60	0.42	0.61	0.38	0.24	0.31	0.27	0.34	0.28

Table 1 – (continued)

Yemeni plateau (Manakhah suite)									
Rock type	Picrite		Basalt						
Lon. (dd)	43.7278	43.7111	43.7347	43.7325	43.7292	43.7247	43.7247	43.7278	
Lat.	15.0881	15.1028	15.0953	15.0881	15.0828	15.0900	15.0900	15.0881	
Sample	YE26	YE29	YE23	YE24	YE25	YE27	YE28	YE31	
SiO ₂ (wt%)	42.22	40.93	37.86	39.20	41.20	38.15	40.31	40.98	
TiO ₂	3.31	3.21	4.16	4.46	3.52	5.07	4.69	3.52	
Al ₂ O ₃	9.98	9.90	9.55	13.25	11.20	11.76	12.10	12.33	
FeO	10.75	11.60	14.09	12.86	12.37	15.41	14.78	10.80	
Fe ₂ O ₃	1.61	1.74	2.11	1.93	1.85	2.31	2.22	1.62	
MnO	0.20	0.19	0.23	0.21	0.20	0.21	0.20	0.17	
MgO	16.21	16.98	12.30	9.46	11.34	6.77	7.96	11.93	
CaO	9.98	10.38	13.93	11.49	13.10	13.00	12.31	12.48	
Na ₂ O	1.74	1.74	1.32	3.81	2.20	2.80	2.26	1.81	
K ₂ O	0.47	0.81	0.20	0.76	0.12	0.92	0.85	1.13	
P ₂ O ₅	0.31	0.43	0.46	0.60	0.47	0.54	0.47	0.71	
LOI	3.21	2.09	3.77	1.98	2.42	3.05	1.84	2.52	
mg#	0.73	0.72	0.61	0.57	0.62	0.44	0.49	0.66	
Ni (ppm)	520	363	214	54	167	1	41	99	
Co	65	59	68	47	63	56	55	55	
Cr	935	673	540	19	556	23	118	279	
V	367	363	521	382	400	522	517	379	
Sc	33	32	54	39	43	53	49	35	
Rb	7	38	3	7	nd	20	14	27	
Sr	671	1071	1065	1106	724	1158	989	1087	
Ba	356	377	478	481	321	500	526	493	
Zr	217	361	281	420	183	349	288	238	
Hf	4.98	8.07	6.14	8.84	4.26	8.02	6.66	5.51	
Nb	56.0	70.5	84.4	85.6	47.8	99.5	75.5	67.9	
Ta	3.74	4.57	5.24	5.46	3.17	6.35	4.84	4.33	
Th	3.65	4.72	5.14	5.51	2.69	6.63	5.20	3.80	
U	1.05	1.50	1.70	1.72	0.88	1.86	1.51	1.15	
Y	19.3	26.1	19.7	33.5	17.6	31.8	27.4	20.8	
La	32.8	46.2	51.2	50.4	30.1	58.7	46.8	40.5	
Ce	72.7	105	112	116	69.6	130	104	91.6	
Pr	8.74	12.7	13.1	14.3	8.52	15.7	12.5	11.0	
Nd	35.9	52.4	51.9	58.5	36.2	64.1	51.4	45.3	
Sm	6.84	9.94	9.17	11.0	6.96	12.0	9.66	8.35	
Eu	2.12	3.01	2.73	3.3	2.13	3.58	2.93	2.57	
Gd	6.66	9.62	8.86	10.8	6.74	11.62	9.49	7.99	
Tb	0.85	1.23	1.06	1.39	0.85	1.45	1.20	0.98	
Dy	4.09	5.81	4.84	6.77	4.08	6.88	5.76	4.57	
Ho	0.71	0.99	0.81	1.19	0.71	1.19	1.00	0.78	
Er	1.91	2.54	2.12	3.21	1.89	3.12	2.67	2.06	
Tm	0.23	0.29	0.24	0.39	0.23	0.37	0.32	0.24	
Yb	1.36	1.71	1.41	2.37	1.34	2.16	1.93	1.43	
Lu	0.18	0.22	0.18	0.32	0.18	0.29	0.26	0.19	

Table 2

Table 2

Rock type <i>n. av. an.</i> sample	MI				GI		
	Picrite				Basalt	Picrite	
	2	1	1	3	1	1	4
	LAL70	LAL73-a	LAL73-b	LAL73-c	LAL42	LAL73-d	LAL73-e
	* (Ol -16%, Fo 87)						
SiO ₂ (wt%)	43.46	47.00	51.09	52.06	52.36	52.45	55.55
TiO ₂	3.12	3.36	5.37	4.48	4.03	2.55	2.22
Al ₂ O ₃	10.87	11.08	12.80	14.33	15.01	16.10	17.97
FeO _{Tot}	15.66	9.96	6.90	7.80	7.81	10.13	6.86
MnO	0.08	0.13	0.07	0.14	0.13	0.21	0.15
MgO	15.42	7.44	5.89	3.33	3.32	1.89	1.13
CaO	8.04	13.76	14.99	14.71	13.28	6.74	4.84
Na ₂ O	1.94	1.55	1.61	2.72	2.81	2.78	4.38
K ₂ O	0.72	1.15	1.48	1.31	1.46	1.64	2.02
P ₂ O ₅	0.50	0.51	0.58	0.83	0.66	0.81	0.95
Cr ₂ O ₃	0.12	0.15	-	-	-	0.02	0.04
NiO	-	0.09	0.25	0.01	0.03	-	-
F	-	0.18	0.22	0.10	0.00	0.16	0.27
Cl	-	0.00	0.03	0.06	0.03	0.05	0.07
Tot	99.93	96.36	101.27	101.88	100.92	95.53	96.44
mg#	0.61	0.60	0.63	0.46	0.46	0.27	0.24
<i>n. av. an.</i>	1		2		1	2	
Sr (ppm)	558		883		838	1178	
Y	19.7		29.9		37.3	36.7	
Zr	202		289		386	458	
Hf	4.71		9.80		9.78	10.55	
Nb	35.2		53.2		43.7	84.3	
Ta	1.53		2.46		2.63	4.87	
La	24.3		38.8		43.6	69.6	
Ce	53.7		89.0		97.0	144.3	
Pr	6.92		10.6		12.72	16.37	
Nd	34.3		45.6		56.6	71.2	
Sm	7.92		10.8		12.84	15.86	
Eu	2.48		3.93		3.83	4.61	
Gd	5.84		9.38		11.14	11.10	
Tb	0.82		1.53		1.76	1.71	
Dy	5.16		7.23		8.40	8.38	
Ho	0.98		1.19		1.60	1.54	
Er	1.90		2.31		3.55	2.79	
Tm	0.20		0.21		0.45	0.52	
Yb	1.39		1.63		3.78	3.37	
Lu	0.18		0.24		0.36	0.44	
Th	1.61		2.78		4.55	8.05	
U	0.43		0.98		1.63	3.03	

* corrected for reheating simulation (after Danyushevsky, 2012)

Table 3

Rock type sample	Ethiopian Plateau (Lalibela suite)																			Yemeni Plateau (Manakhah suite)		
	Picrite								Basalt											Basalt		
	LAL41		LAL45		LAL70	LAL73		LAL31		LAL 42		LAL43		LAL44		LAL65		LAL69		YE28		
Cpx crystal n.	1	5	2	5	3	5	3	2	4	1-2	5	1	2	1	2	2	4	3	4	1	2	3
<i>n. av. an.</i>	2	2	2	2	3	3	2	3	4	3	6	4	2	2	2	2	3	2	4	3	3	3
Rb (<i>ppm</i>)	0.01	0.01	0.01	0.02	0.47	0.00	0.00	0.01	1.54	0.04	0.11	0.01	0.02	0.01	0.01	0.40	0.17	0.12	0.03	0.01	0.02	0.02
Sr	40.4	35.1	58.7	75.6	66.6	57.9	81.6	52.1	52.8	48.2	48.0	69.2	65.9	69.0	56.6	85.8	87.5	86.8	75.2	117	109	92.1
Ti	6978	4062	6864	10012	9374	7632	7260	7400	5634	7398	7389	7494	8153	7128	5015	12083	9733	10869	8981	11286	11271	10231
Y	9.20	5.13	8.70	13.7	16.8	9.16	7.45	14.3	8.67	13.4	11.4	8.60	8.06	7.82	3.69	19.09	19.17	16.1	12.9	24.1	18.2	13.8
Zr	20.8	7.11	21.3	50.5	49.4	30.0	22.9	34.5	28.6	30.1	28.4	23.5	23.9	19.5	7.06	70.71	63.84	66.5	45.4	125	86.3	70.7
Hf	1.10	0.37	1.18	2.46	2.17	1.74	0.92	1.67	1.11	1.33	1.31	1.26	1.06	1.19	0.28	3.39	2.91	3.45	2.29	5.20	3.82	3.28
Nb	0.06	0.01	0.09	0.21	0.32	0.29	0.39	0.15	0.28	0.41	0.10	0.07	0.20	0.07	0.03	0.31	0.26	0.31	0.13	0.96	0.59	0.43
Ta	0.01	0.00	0.01	0.03	0.02	0.01	0.02	0.01	0.03	0.02	0.01	0.01	0.00	0.01	0.00	0.05	0.03	0.04	0.01	0.17	0.11	0.08
La	1.24	0.71	1.71	3.26	3.41	2.20	2.02	2.08	2.43	1.96	1.82	1.68	1.68	1.47	0.79	3.73	3.36	3.93	2.97	11.4	5.94	5.56
Ce	5.19	3.21	7.41	13.2	13.8	7.56	7.77	10.3	7.17	7.89	8.06	7.91	8.13	6.24	4.14	15.14	14.48	16.7	11.9	43.6	25.2	22.4
Pr	1.13	0.64	1.37	2.44	2.67	1.53	1.30	2.08	1.18	1.46	1.49	1.58	1.52	1.34	0.75	3.04	2.86	3.12	2.40	7.54	4.43	3.82
Nd	6.90	4.04	8.2	13.9	17.6	7.89	7.79	11.0	6.83	10.4	9.79	8.64	9.63	8.11	4.43	18.7	18.0	19.1	14.5	39.5	25.7	20.2
Sm	2.41	1.70	2.85	5.18	4.82	3.12	2.42	4.05	2.32	4.24	3.03	3.34	4.10	2.94	1.91	6.13	5.99	5.86	4.54	10.1	6.85	5.70
Eu	0.85	0.54	1.05	1.51	1.62	0.83	0.88	1.37	0.72	1.13	1.22	1.16	1.17	0.92	0.57	2.21	2.05	2.09	1.58	3.16	2.57	1.84
Gd	2.72	1.64	2.82	4.15	4.98	2.70	2.63	4.22	2.34	3.87	2.96	3.06	3.23	2.32	1.46	5.74	6.15	5.61	4.10	8.26	5.72	4.51
Tb	0.40	0.22	0.43	0.63	0.73	0.39	0.35	0.71	0.34	0.55	0.47	0.44	0.43	0.36	0.24	0.86	0.93	0.79	0.59	1.16	0.92	0.66
Dy	2.42	1.48	2.16	3.30	4.42	2.10	2.08	3.80	1.99	3.41	2.78	2.02	2.28	2.12	1.22	5.01	4.89	3.75	3.42	6.21	5.03	3.43
Ho	0.39	0.24	0.38	0.60	0.78	0.41	0.29	0.70	0.38	0.55	0.49	0.37	0.39	0.33	0.19	0.80	0.83	0.67	0.50	1.04	0.78	0.58
Er	0.80	0.69	0.88	1.35	1.40	0.59	0.68	1.49	0.83	1.42	0.98	0.75	0.90	0.78	0.29	1.91	1.90	1.59	1.20	2.74	1.80	1.39
Tm	0.11	0.07	0.10	0.16	0.20	0.06	0.11	0.18	0.10	0.16	0.13	0.10	0.12	0.07	0.03	0.18	0.20	0.15	0.15	0.32	0.22	0.17
Yb	0.73	0.36	0.70	0.81	1.27	0.48	0.55	1.01	0.67	1.19	0.96	0.60	0.59	0.32	0.20	1.26	1.41	1.15	0.97	1.84	1.48	1.08
Lu	0.08	0.05	0.07	0.11	0.19	0.05	0.07	0.14	0.07	0.12	0.12	0.09	0.10	0.06	0.03	0.15	0.14	0.11	0.12	0.23	0.20	0.11
Th	0.01	0.01	0.01	0.02	0.08	0.03	0.01	0.01	0.32	0.09	0.03	0.01	0.09	0.01	0.00	0.04	0.03	0.02	0.01	0.10	0.03	0.03
U	0.00	0.00	0.00	0.00	0.02	0.01	0.02	0.00	0.23	0.03	0.02	0.00	0.01	0.00	0.00	0.01	0.01	0.01	0.00	0.01	0.00	0.01
Pb	0.04	0.17	0.09	0.07	0.17	0.18	0.25	0.19	4.91	0.19	0.52	0.16	0.14	0.13	0.22	0.03	0.12	0.09	0.04	0.14	0.09	0.10

Table 4

Affinity	sample		$^{87}\text{Sr}/^{86}\text{Sr}$	$^{87}\text{Sr}/^{86}\text{Sr}_i$	$^{143}\text{Nd}/^{144}\text{Nd}$	$^{143}\text{Nd}/^{144}\text{Nd}_i$	$^{206}\text{Pb}/^{204}\text{Pb}$	$^{206}\text{Pb}/^{204}\text{Pb}_i$	$^{207}\text{Pb}/^{204}\text{Pb}$	$^{207}\text{Pb}/^{204}\text{Pb}_i$	$^{208}\text{Pb}/^{204}\text{Pb}$	$^{208}\text{Pb}/^{204}\text{Pb}_i$
Ethiopian Plateau												
LT-Bas	SIM15	Bulk Rock	0.70367	0.70367	0.51288	0.51285	18.44	18.42	15.55	15.55	37.91	37.89
LT-Bas	SIM17	Bulk Rock	0.70373	0.70373	0.51280	0.51277	18.22	18.19	15.53	15.53	37.67	37.63
LT-Bas	ADG3	Bulk Rock	0.70329	0.70329	0.51306	0.51303	18.31	18.30	15.51	15.51	37.97	37.95
LT-Bas	ADG6	Bulk Rock	0.70320	0.70320	0.51298	0.51295	18.32	18.30	15.52	15.52	37.92	37.91
HT1-Bas	LAL32	Bulk Rock	0.70368	0.70366	0.51302	0.51300	18.71	18.63	15.56	15.56	38.50	38.45
HT1-Bas	LAL77	Bulk Rock	0.70412	0.70408	0.51289	0.51286	18.59	18.48	15.52	15.52	38.13	38.05
HT1-Bas	BLN2	Bulk Rock	0.70429	0.70428	0.51292	0.51290	18.86	18.79	15.59	15.59	38.47	38.39
HT1-Bas	BLN4	Bulk Rock	0.70371	0.70369	0.51294	0.51291	18.57	18.51	15.51	15.51	38.37	38.29
HT2-Bas	LAL14	Bulk Rock	0.70424	0.70420	0.51295	0.51293	19.26	19.11	15.56	15.55	38.76	38.58
HT2-Bas	LAL31	Bulk Rock	0.70415	0.70408	0.51299	0.51296	18.95	18.72	15.58	15.57	38.76	38.58
HT2-Bas	LAL42	Bulk Rock	0.70437	0.70432	0.51296	0.51293	18.96	18.77	15.57	15.56	38.41	38.27
		Cpx	0.70436	0.70436	0.51299	0.51294	18.83	18.81	15.59	15.59	38.45	38.43
HT2-Bas	LAL43	Bulk Rock	0.70406	0.70402	0.51303	0.51300	19.12	18.92	15.59	15.58	38.93	38.72
HT2-Bas	LAL44	Bulk Rock	0.70417	0.70412	0.51295	0.51292	19.02	18.84	15.55	15.54	38.87	38.72
		Cpx	0.70467	0.70412	0.51300	0.51295	18.45	18.45	15.40	15.40	38.23	38.23
HT2-Bas	LAL65	Bulk Rock	0.70433	0.70427	0.51301	0.51299	19.03	18.90	15.58	15.57	38.96	38.87
		Cpx	0.70429	0.70429	0.51301	0.51297	18.81	18.77	15.60	15.60	38.77	38.73
HT2-Bas	LAL68	Bulk Rock	0.70418	0.70415	0.51294	0.51291	19.07	18.88	15.58	15.57	38.96	38.82
HT2-Bas	LAL69	Bulk Rock	0.70412	0.70404	0.51302	0.51230	19.08	18.97	15.58	15.58	39.13	39.04
		Cpx	0.70412	0.70412	0.51303	0.51299	18.96	18.94	15.64	15.64	39.06	39.04
HT2-Pi	LAL33	Bulk Rock	0.70409	0.70405	0.51300	0.51297	18.94	18.82	15.56	15.55	38.71	38.61
HT2-Pi	LAL41	Bulk Rock	0.70417	0.70408	0.51296	0.51294	18.95	18.83	15.59	15.58	38.76	38.65
HT2-Pi	LAL45	Bulk Rock	0.70401	0.70394	0.51298	0.51296	19.05	18.93	15.58	15.58	39.13	38.72
HT2-Pi	LAL47	Bulk Rock	0.70418	0.70412	0.51296	0.51293	19.08	18.97	15.58	15.58	38.84	38.76
HT2-Pi	LAL70	Bulk Rock	0.70404	0.70399	0.51290	0.51288	19.07	18.97	15.63	15.63	38.99	38.93
		Cpx	0.70454	0.70453	0.51297	0.51294	18.51	18.48	15.43	15.42	38.30	38.26
HT2-Pi	LAL73	Bulk Rock	0.70414	0.70407	0.51294	0.51292	18.99	18.88	15.55	15.55	38.89	38.83
		Cpx	0.70439	0.70439	0.51301	0.51297	18.61	18.59	15.54	15.54	38.56	38.55
Yemeni Plateau												
HT2-Bas	YE23	Bulk Rock	0.70363	0.70363	0.51290	0.51290	19.41	19.24	15.61	15.60	39.46	39.29
HT2-Bas	YE24	Bulk Rock	0.70359	0.70358	0.51299	0.51297	18.91	18.74	15.55	15.54	38.89	38.71
HT2-Bas	YE25	Bulk Rock	0.70362	0.70362	0.51290	0.51288	19.22	19.13	15.55	15.55	39.21	39.12
HT2-Bas	YE27	Bulk Rock	0.70371	0.70368	0.51288	0.51285	19.26	19.07	15.55	15.54	39.26	39.04
HT2-Bas	YE28	Bulk Rock	0.70370	0.70368	0.51289	0.51287	18.96	18.81	15.57	15.56	39.01	38.84
HT2-Bas	YE31	Bulk Rock	0.70368	0.70365	0.51293	0.51291	19.38	19.26	15.61	15.61	39.39	39.26
HT2-Pi	YE26	Bulk Rock	0.70372	0.70371	0.51289	0.51287	19.00	18.89	15.60	15.60	39.13	39.01
HT2-Pi	YE29	Bulk Rock	0.70367	0.70362	0.51295	0.51293	19.00	18.85	15.57	15.56	38.95	38.80

Table 5

Affinity	sample	$^3\text{He}/^4\text{He}$ (R/R _A)		
Ethiopian Plateau				
HT2-Bas	LAL31	17.8	±	0.7
HT2-Bas	LAL42	12.9	±	0.7
HT2-Bas	LAL43	19.1	±	1.3
HT2-Bas	LAL44	15.6	±	0.4
HT2-Pi	LAL6	10.4	±	0.9
HT2-Pi	LAL11	10.6	±	0.6
HT2-Pi	LAL33	15.7	±	1.0
HT2-Pi	LAL41	16.8	±	0.5
HT2-Pi	LAL47	18.3	±	0.6
HT2-Pi	LAL70	3.7	±	0.8
HT2-Pi	LAL73	9.6	±	1.0

Table 6

Affinity	sample	mineral	$\delta^{18}\text{O}$ (‰)		
Ethiopian Plateau					
HT2-Bas	LAL42	Olivine	6.12	±	0.08
HT2-Bas	LAL42	Cpx	5.82	±	0.09
HT2-Bas	LAL70	Olivine	5.21	±	0.17
HT2-Bas	LAL70	Cpx	5.82	±	0.22
HT2-Pi	LAL73	Olivine	6.62	±	0.21
HT2-Pi	LAL11	Cpx	6.90	±	0.08